## Communications Research <br> Centre

# APPLICATION OF GEOLOCATION TECHNIQUES USING SATELLITES IN GEOSTATIONARY ORBIT (U) 

by

Mario Caron<br>(Communications Technologies Research Branch)

COMMUNICATIONS RESEARCH
CENTRE

## DEPARTMENT OF COMMUNICATIONS CANADA

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#### Abstract

This report looks at techniques to perform the localization of SARSAT 406 MHz distress beacons using geostationary satellites. After reviewing the characteristics of the distress beacon transmitters, the main specifications of the GOES satellites are described as they can be assumed as typical geostationary satellites for this application. A brief review of the geolocation techniques is presented and an in-depth analysis is conducted on the time difference of arrival technique. The analysis covers the impact of the Earth's flatness, channel impairments and satellite geometry . The delay estimator performance is expressed as a function of signal-to-noise ratio and circular error probability. The overall system performance is directly related to the performance of the delay estimator.

Results of both theoretical analysis and computer simulations show that the accuracy of the position derived from the time difference of arrival technique is limited by (1) the low SNR of the distress beacons when relayed via geostationary satellites and, (2) by the low transmission rate (and thus bandwidth) of the beacon signal creating "broad" autocorrelation peaks and making the delay estimation process difficult. In the worst case $\mathrm{C} / \mathrm{No}$ of $30 \mathrm{~dB}-\mathrm{Hz}$ and with three satellites spaced at $30^{\circ}$, we can say with a $90 \%$ confidence level that the circular error probability is between 3.3 and 4.9 km for a beacon in Canada. This is equivalent to an error not exceeding between 8 and 11.8 km for $95 \%$ of the time which is comparable to the current SARSAT polar orbit system based on a maximum error of 5 km for $90 \%$ of the time.


## résumé

Ce rapport examine les techniques permettant d'effectuer la localisation de radiobalises de détresse SARSAT à 406 MHz en utilisant des satellites géostationnaires. Après une revue des caractéristiques des émetteurs des radiobalises de détresse, les spécifications principales des satellites GOES sont décrites puisqu'ils peuvent être considéré comme des satellites géostationnaires typiques pour cette application. Une brève revue des techniques de géolocation est présentée et une analyse en profondeur de la technique des différences de temps est effectuée. L'analyse couvre l'impact de l'applatissement de la Terre, les détériorations du canal et la géométrie des satellites. La performance de l'estimateur de délai est exprimée en fonction du rapport signal-àbruit et de la probabilité d'erreur circulaire. La performance du système global est directement reliée à la performance de l'estimateur de délai.
$\underline{L}$

Les résultats de l'analyse théorique et des simulations sur ordinateur démontrent que la précision de la position déduite à partir de la technique des différences de temps est limitée par (1) le faible rapport signal-à-bruit des radiobalises de détresse lorsque retransmises par satellites géostationnaires et, (2) par le faible débit de transmission (et ainsi de largeur de bande) du signal de la radiobalise créant de "larges" crêtes d'autocorrelation et rendant le processus d'estimation du délai difficile. Dans le pire cas d'un rapport C/No de $30 \mathrm{~dB}-\mathrm{Hz}$ et avec trois satellites espacés de $30^{\circ}$, nous pouvons affirmer avec un degré de confiance de $90 \%$ que la probabilité d'erreur circulaire est entre 3.3 et 4.9 km pour une radiobalise au Canada. Ceci est équivalent à une erreur ne dépassant pas entre 8 et 11.8 km pour $95 \%$ du temps qui est comparable au système actuel basé sur une erreur maximale de 5 km pour $90 \%$ du temps.

## EXECUTIVE SUMMARY

The current Geostationary Operational Environmental Satellites (GOES's) are equipped with a special repeater allowing 406 MHz SARSAT distress signals to be relayed to the Mission Control Centres (MCC's). When a distress beacon is activated and is within the field of view of the satellite, this special repeater allows for an instantaneous distress alerting. This is particularly important for distress events occurring around the equator where the sole use of the SARSAT/ COSPAS low earth orbit satellites can result in up to one hour delay before the distress event is reported. Geostationary satellites have a continuous coverage of this area and thus can improve significantly the overall response time of the system. The SARSAT/COSPAS system currently relies on the Doppler frequency shift on the beacon signal to determine the beacon position. Geostationary satellites being quasi-motion free relative to the Earth, there is very little Doppler frequency shift that can be used to locate the distress beacons and an alternative technique must be used. This report investigates such alternative techniques.

With geostationary satellites, it is found that only the time difference of arrival technique can be promising. At least three satellites are required for an unambiguous positioning. The analysis shows that the Earth's flatness and the channel (ionosphere and troposphere) delay has little impact on the overall position accuracy of the system. Indeed, it is shown that the accuracy is more limited by the low signal-to-noise ratio of the beacon signal once relayed via a geostationary satellite and of its low bandwidth (or data rate) which creates relatively broad autocorrelation peaks and make the estimation of very precise delay difficult. The theoretical analysis and the computer simulations show that an accuracy comparable to the current SARSAT polar orbit system can be expected i.e. an error less than 5 km in $90 \%$ of the time.

The position estimator can be based on a personal computer equipped with special digital boards to perform the signal processing. The implementation proposed in this report assumed that the position estimation processor is an add-on to the current GOES processor. Implementation should however be deferred until sufficient GOES or other host satellites become available. More advanced techniques to perform the delay estimation could be investigated in the mean time.

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### 1.0 INTRODUCTION

The work reported here has been funded by the EHF Satcom division of the Defence Research Establishment of Ottawa (DREO). Their interest in the application of geolocation techniques using geostationary satellites is twofold. First it provides a tool to locate sources of interference so that, for instance, the satellite antenna/nulling system can generate a null on the interference source. Second, it is a tool to assess the probability that covert terminals can be detected and located, or the success of the electronic counter-countermeasure (ECCM) technique(s) used by the terminal. In general, these signal sources are random, bursty, and have poor signal-tonoise ratios.

The 406 MHz emergency beacons transponded through the Geostationary Operational Environmental Satellite (GOES) satellites have characteristics similar to the signal sources of interest. It is the purpose of this report to investigate the feasibility of developing a system to extract location information from 406 MHz distress signals that have been transmitted by repeaters on GOES or other similar geostationary satellites. Because geostationary satellites offer continuous coverage, they have been considered in various systems for radio position determination applications [1-11]. The major benefit for the search and rescue (SAR) system lies in the possibility of improving the response time of the current system by using geostationary satellites. Indeed, the use of geostationary satellites provides a quasi-instantaneous position reporting for beacons in their field of view while some distress alerts in the current COSPASSARSAT system may take up to two hours to be detected and located. This is partly due to the ambiguity of the Doppler positioning technique which requires a second satellite pass to resolve and partly due to the satellite orbit which makes the satellite revisit time relatively long for terminals close to the equator. The use of a geostationary satellite would therefore be particularly useful to remove these latter delays where the geostationary satellite visibility is continuous.

Sections 2.0 and 3.0 of this report review briefly the characteristics of 406 MHz emergency beacons and GOES repeaters respectively. Section 4.0 describes a number of basic geolocation techniques and discusses their applicability to this project. Section 5.0 gives the basic principle of the time difference of arrival (TDOA) technique while Section 6.0 introduces the perturbations to the TDOA principle. Section 7.0 deals exclusively with the design of the time difference estimator. Section 8.0 combines the effects of all perturbations and derives an expression for the overall positioning accuracy of a proposed distress beacon receiver. Section 9.0 describes the circular error probability of the system and Section 10.0 summarizes the receiver structure. Section 11.0 concludes this report.

### 2.0406 MHZ EMERGENCY BEACON CHARACTERISTICS

The detailed specifications of the 406 MHz emergency beacon transmitters are given in [12] and this section briefly highlights some of the key signal characteristics.

The beacon transmits a digitally phase modulated carrier burst. The format of the burst is determined by the message type i.e. short (standard) or long (optional). Figure 2.0-1 along with Table 2.0-1 shows the signal format. Basically, the burst duration is either 440 ms (short) or 520 ms (long) and it is repeated every T seconds where T is a random period uniformly distributed between 47.5 and 52.5 seconds. The signal comprises a 160 ms carrier preamble and a 24 -bit synchronization word. The protected data field is 61 bits long and the error correcting code is a $\mathrm{BCH}(127,106)$ code shortened to $(82,61)$. The following 6 bits are reserved for national use or emergency codes while the last 32 bits are the long message (optional) data. The optional message content is not strictly defined but it is recommended in [12] to maritime users to include the course, speed and time of activation while other users are recommended to include the beacon position derived from an auxiliary radio location system incorporated into the beacon. Of course this study is limited to those beacons which are not equipped with such auxiliary location systems.

The phase modulation is such that a residual carrier is present to ease the signal detection and demodulation. The initial frequency accuracy is $\pm 2 \mathrm{kHz}$ and the frequency offset does not exceed 5 kHz after 5 years. The power output is nominally 5 watts with an antenna gain between -3 dBi and 4 dBi over $90 \%$ of the hemispherical pattern.


Figure 2.0-1 General 406 MHz Message Format

| Parameter | Value |
| :---: | :---: |
| RF Signal <br> Carrier frequency (initial) <br> Carrier frequency (prior to <br> end of useful life) <br> Frequency stability <br> Short term ( 100 ms ) <br> Medium term <br> Mean slope <br> Residual noise <br> Power output <br> Spurious emissions <br> Data encoding <br> Modulation <br> Modulation rise and fall times <br> Digital Message <br> Repetition rate <br> Transmission Time <br> CW preamble <br> Digital message <br> Bit rate <br> Bit synchronization <br> Frame synchronization <br> Continuous emission: failure mode | $406.025 \pm 0.002 \mathrm{MHz}$ <br> $406.025 \pm 0.005 \mathrm{MHz}$ <br> $2 \times 10^{-9}$ $1 \times 10^{-9} / \mathrm{min}$ $3 \times 10^{-9}$ <br> $5 \mathrm{~W} \pm 2 \mathrm{~dB}$ into 50 ohm <br> load with VSWR < 1.25:1 <br> 50 dB below 5 W in 5 MHz <br> bandwidth; carrier harmonics 30 dB <br> below 5 W <br> Bi-phase L <br> Phase modulation <br> $1.1 \pm 0.1$ radians peak must be between $50 \mu \mathrm{~s}$ and $250 \mu \mathrm{~s}$ <br> Random. Uniform Distribution <br> Between 47.5 and 52.5 s <br> $440 \mathrm{~ms} \pm 1 \%$ or $520 \mathrm{~ms} \pm 1 \%$ <br> $160 \mathrm{~ms} \pm 1 \%$ <br> $280 \mathrm{~ms} \pm 1 \%$ or $360 \mathrm{~ms} \pm 1 \%$ <br> $400 \mathrm{bps} \pm 1 \%$ <br> 15 "ones" <br> 000101111 <br> Transmission shall not exceed 45 s |

Table 2.0-1 406 MHz Beacon Signal Characteristics

### 3.0 GOES REPEATER CHARACTERISTICS

The latest Geostationary Operational Environmental Satellite (GOES H) is equipped with a frequency translating repeater which receives signals over a 100 kHz band centered at 406.05 MHz and re-transmits the translated and inverted band at a center frequency of 1698.65 MHz . The 406 MHz repeater on the current GOES H satellite shares circuitry with another GOES subsystem, the Data Collection Platform (DCP) repeater. Both subsystems have different operating bands but a common AGC regulates the total output power of the combined DCP and 406 MHz repeater. Note that the GOES satellites subsequent to GOES H will not have this common AGC and the translation frequency band will be slightly different.

The satellite antenna receives right hand circularly polarized signals and transmits linearly polarized signals. The antenna pattern is hemispheric in both receive and transmit. The repeater gain variation over $406.05 \mathrm{MHz} \pm 40 \mathrm{kHz}$ does not exceed 2 dB peak-to-peak.

The minimum satellite $\mathrm{G} / \mathrm{T}$ is $-22 \mathrm{~dB} / \mathrm{K}$ and the EIRP is +30 dBm shared by 406 MHz distress signals when no transmission occurs on the DCP repeater. During DCP transmission, the 406 MHz beacons available EIRP drops by approximately 6 dB based on experimental results.

Link budgets for a $5^{\circ}$ elevation angle and 8 simultaneous distress signals given in [14] for a GOES satellite at either $75^{\circ} \mathrm{W}$ or $135^{\circ} \mathrm{W}$ show that the minimum expected $\mathrm{C} / \mathrm{No}$ is in the order of $30 \mathrm{~dB}-\mathrm{Hz}$.

The GOES satellites are maintained in their nominal orbital position within approximately $\pm 0.1^{\circ}$ in both the East-West and North-South axis.

Note that the current GOES satellite has a CW pilot that can be used for frequency tracking. In addition, time is disseminated at 468.8 MHz . This time code generated by atomic clocks is repeated every 30 seconds. Its accuracy is not currently known but commercial receivers offered by Kinemerrics/Truetime in California give an accuracy of 0.5 ms .

### 4.0 GEOLOCATION TECHNIOUES

Geolocation refers to the determination of the location of a vehicle, object or person on earth. The location of signal sources is usually based on techniques which rely on one or a combination of frequency, time and spatial information. It is difficult to describe various geolocation techniques that can be based on a combination of this information. The approach taken here is to identify how the information can be used when a limited number of satellites is available. This is realistic and limits the study to practical cases. This section describes techniques applicable to one, two and more than two satellite systems in a general sense. The time difference of arrival technique is discussed in detail in the following sections.

### 4.1 ONE-SATELLITE SYSTEMS

In general, the geolocation systems based on a single satellite require special circuits to be built into the spacecraft. These circuits are, for instance, an interferometer, a steering antenna with an energy detector scanning the Earth as in [15] or a special antenna mounted on a spacecraft with a specific 3 -dimensional motion as in [16,17]. Clearly these special circuits are not available on GOES satellites.

A technique that does not require special circuits is based on the residual velocity of the nearly geostationary satellites. Indeed geostationary satellites are never fully stationary because of the orbit perturbations such as the changes with time of the gravitational pull of the moon and to a lesser extent of the sun [18]. The geostationary satellites are usually maintained in their nominal positions with the East-West and North-South station keeping controls.

The slow geostationary satellite motion creates a small Doppler frequency shift. To have an idea of the order of magnitude of this Doppler shift, Slabinsky [18] gives an example with a $\pm 0.25^{\circ}$ East-West and $\pm 0.1^{\circ}$ North-South maximum position error and a satellite eccentricity $\varepsilon$ for a station located at $40^{\circ}$ North latitude and $50^{\circ}$ West of the satellite meridian. The worst case satellite to ground station radial velocity is given by $(0.56+3100 \varepsilon) \mathrm{m} / \mathrm{s}$. The amplitude of range variation at the ground station is $11.8 \mathrm{~km} /$ day about a mean value which changes less than $2.8 \mathrm{~km} /$ day. Although the GOES satellite eccentricity is not known, typical values for geostationary satellites are between $4.7 \times 10^{-5}$ and $3.6 \times 10^{-4}$. With an eccentricity of $3.6 \times 10^{-4}$, it gives a maximum residual velocity of $1.68 \mathrm{~m} / \mathrm{s}$ which corresponds to a maximum Doppler frequency shift of 2.27 Hz for a carrier at 406 MHz . This is obviously too small to be used in a positioning system as it would require very good short term oscillator stability on-board the
spacecraft and/or complex circuits on the ground station to measure accurately a fraction of this Doppler frequency shift.

In a similar way, the variation of time delay with time could be monitored to derive the source location. Although the knowledge of the time delay variation with time does not give us an unambiguous source position, it delineates an area to search. Based on the example above, the range variation is bounded to $11.8 \mathrm{~km} /$ day or $492 \mathrm{~m} / \mathrm{hour}$. This corresponds to a time delay variation of $1.6 \mu \mathrm{~s} /$ hour. Given that this is a maximum rate of change which may be difficult to measure over such a long period of time and given that the goal of this system is to provide an "instantaneous" response, this information of the time delay variation is of little use.

### 4.2 TWO-SATELLITE SYSTEMS

The use of two satellites provides spatial diversity that can be exploited in a number of ways. First, the two satellite system can be seen as a phased array antenna with two elements and interferometer direction finding (DF) techniques can in principle be used. An interferometer relies on the phase or time difference of arrival between two or more antennas. It provides an unambiguous position location provided that the spacing between the two antennas is less than half a wavelength. When the distance exceeds half a wavelength, some techniques usually involving additional antennas must be used to resolve the ambiguity. Because the wavelength at 406 MHz is 0.73 m and the two satellites will likely be many degrees apart to have a maximum coverage area, this technique is not applicable.

An alternate approach would be to measure the time of arrival of the signals from each satellite. This technique has become very popular over the last few years. Several positioning systems based on this principle are currently available or under study [1-11, 19-22]. Although the type of signals and required equipment vary from one system to the other, they are all used to determine the position of the source based on the absolute time of arrival of the signal from two or more satellites or on the time difference of arrival between these signals.

In a two-way system where the source transmits only after being polled by the master station, the time delay is easily obtained from the time the signal left the master station and the time the reply is received. Signals received from two or more satellites are then used to estimate the source location. In a one-way system where the system can transmit at any time, there is no common time base and the position must be determined from the time difference between three or more satellites i.e. given the propagation delay differences between three or more satellites, it is
possible to determine where the source is located. For the location of the 406 MHz distress signals, only the one way technique is applicable and with only two satellites, the propagation delay difference from two satellites is insufficient to provide an unambiguous position location even if the satellite positions are known exactly. A minimum of three satellites is required for the one-way technique.

### 4.3 MORE THAN TWO-SATELLITE SYSTEMS

Basically, the techniques applicable to the two-satellite systems can also be extended to three or more satellites systems. As mentioned before if three satellites are used, then the beacon location can be determined using the time difference of arrival of the signals between each satellite and an estimate of the beacon altitude. The use of a fourth satellite makes the estimate of the altitude unnecessary. For the interferometer technique, the use of more satellites reduces the number of ambiguities but it is unlikely that it will be able to provide a good position location with practical satellite spacing.

### 5.0 TIME DIFFERENCE OF ARRIVAL BASIC PRINCIPLE

The differential time of arrival technique was described briefly in the previous section. It is worthwhile to review its theory of operation in more detail. Figure 5.0-1 illustrates the scenario. When the beacon is activated, a 440 or 520 ms burst message is transmitted every 50 seconds nominally. The signal is received by two or more satellites, up-converted in frequency to the vicinity of 1.7 GHz and then relayed to a master station. The master station logs at which time the signal from each satellite was received and performs some processing (to be discussed below) to derive the beacon location. Because the beacon can be activated at any time, it is not possible to estimate the full time delay between the transmission from the beacon and the reception at the master station i.e. $\mathrm{T}_{1}+\mathrm{T}_{1}{ }^{\prime}$ or $\mathrm{T}_{2}+\mathrm{T}_{2}{ }^{\prime}$. This lead us in the previous section to consider the differential techniques. Referring to Figure 5.0-1 and assuming that the master station range to each satellite is known, then it is possible to get an estimate of the time difference of arrival given by : $\Delta T=T_{2}-T_{1}$. This time delay corresponds to a differential range $\Delta R=R_{2}-R_{1}$. Given this differential range and the satellite locations, it is possible to determine a line of position on the earth where the beacon transmitter could be located. To illustrate this, let us consider a perfectly spherical earth with satellites located above the equator in geostationary orbits at longitudes $\lambda_{0 \mathrm{i}}$ where $\mathrm{i}=1$ or 2 . Considering the geometry shown in Figure $5.0-2$, where the ground station is assumed to be located at latitude $\varphi$ and longitude $\lambda$, the range to the satellite \#i is given by :

$$
\begin{equation*}
R_{i}=R_{e} \frac{\sin \beta_{i}}{\sin \gamma_{i}} \quad \text { meters, } \quad i=1,2 \tag{5.0-1a}
\end{equation*}
$$

where $\mathrm{R}_{\mathrm{e}}=$ equivalent radius of the earth $=6,370,997$ meters

$$
\begin{align*}
& \beta_{\mathrm{i}}=\cos ^{-1}\left[\cos \varphi \cos \left(\lambda-\lambda_{0 \mathrm{i}}\right)\right]  \tag{5.0-1b}\\
& \gamma_{\mathrm{i}}=\tan ^{-1}\left(\frac{\sin \beta_{\mathrm{i}}}{\mathrm{R}_{\mathrm{S}} / \mathrm{R}_{\mathrm{e}}-\cos \beta_{\mathrm{i}}}\right) \tag{5.0-1c}
\end{align*}
$$

$\mathrm{R}_{\mathrm{S}}=$ satellite orbit radius $=42,157,197$ meters


Figure 5.0-1 Overall Block Diagram of Distress Beacon Position Location System

satelicemtaic sphere

Figure 5.0-2 Basic Geometry for the Computation of Satellite Ranges on a Spherical Earth (from [23])

Note that latitude is positive north and longitude is positive east in all of the above equations and in the remainder of this report. If we define $\Delta \lambda=\lambda-\lambda_{0 i}$ and re-arrange, we obtain:

$$
\begin{equation*}
\mathrm{R}_{\mathrm{i}}=\sqrt{\mathrm{R}_{\mathrm{s}}{ }^{2}-2 \mathrm{R}_{\mathrm{s}} \mathrm{R}_{\mathrm{e}} \cos \varphi \cos \left(\Delta \lambda_{\mathrm{i}}\right)+\mathrm{R}_{\mathrm{e}}^{2}} \tag{5.0-2}
\end{equation*}
$$

and

$$
\Delta R_{i j}=R_{j}-R_{i}
$$

A computer program based on existing software for satellite spherical geometry calculation has been developed to implement these equations. By trial and error, it finds the station longitude that generates a given differential range $\Delta \mathrm{R} \pm \delta$ for given satellite locations ( $\lambda_{01}$ and $\lambda_{02}$ ). Figure 5.0-3 illustrates graphically the results for $\lambda_{01}=135^{\circ} \mathrm{W}, \lambda_{02}=75^{\circ} \mathrm{W}, \Delta \mathrm{R}=200 \mathrm{~km}$ and $\delta=1$ meter.

To resolve the positioning ambiguity obtained from two satellites, a third satellite must be used to generate at least another positioning line. The intersection of two such lines gives the beacon transmitter location. Figure 5.0-4 illustrates this process for a transmitter located at Ottawa $\left(45.35^{\circ} \mathrm{N}, 75.90 \mathrm{~W}\right.$ ) and three satellites located respectively at $75^{\circ} \mathrm{W}, 105^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$. The differential ranges were :

$$
\begin{aligned}
& \Delta \mathrm{R}_{12}=622,078.35 \mathrm{~m} \\
& \Delta \mathrm{R}_{23}=1,724,172.32 \mathrm{~m} \\
& \Delta \mathrm{R}_{13}=2,346,250.67 \mathrm{~m}
\end{aligned}
$$

Accordingly the accuracy to which we can locate the terminal is limited by the computation accuracy for this ideal case. In the following sections, we introduce some perturbations to this model which will add some errors to the estimated differential ranges. It is interesting to note that in absence of errors we should always have :

$$
\Delta R_{12}+\Delta R_{23}-\Delta R_{13}=0
$$

This simple relationship could be used in an operational system as a measure of the quality of the estimate.


Figure 5.0-3 Position Line Generated with Satellites at $75^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$ and a Differential Range of 200 km .


Figure 5.0-4 Position Lines For Three Satellites (750 W, $105^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$ ). The Station is at the Intersection of the Lines i.e. in Ottawa.

### 6.0 TDOA PERTURBATIONS

### 6.1 NON-SPHERICAL EARTH

In Section 5.0 we assumed a spherical earth. In practice, the earth is better represented as an oblate spheroid ${ }^{1}$ with equatorial radius $\mathbf{a}$, eccentricity $\mathbf{e}$ and semi-major axis $\mathbf{b}$ (see Figure 6.11). The semi-major axis $\mathbf{a}$ and the semi-minor axis $\mathbf{b}$ are related by :

$$
\begin{equation*}
b=a \sqrt{\left(1-e^{2}\right)} \tag{6.1-1}
\end{equation*}
$$

When using such an earth model, it is important to make the distinction between geocentric and geodetic latitudes. Figure 6.1-1 shows the point $P$ with geocentric latitude $\theta$ and geodetic latitude $\phi$. The local height above the surface of the earth is given by the length of the normal to the oblate spheroid. The geodetic $(\varphi)$ and geocentric $(\theta)$ latitudes are related by the following :

$$
\begin{equation*}
\tan \theta=\left(1-\mathbf{e}^{2}\right) \tan \varphi \tag{6.1-2}
\end{equation*}
$$

To find the point on the earth which generated a given differential range from two given satellites, an approach different than the one in the previous section must be used. Indeed, it is easier to work in what is called the earth centered earth fixed (ECEF) ( $x, y, z$ ) coordinate system than with spherical coordinates. In the ECEF coordinate system, the origin is at the earth center of mass, the x -axis goes through the Greenwich meridian at the equator, the z -axis is the polar axis and the $y$-axis completes the right-hand coordinate system.

Using the notation of Figure 6.1-1 and assuming that the beacon transmitter is located on longitude $\lambda$. The ECEF coordinates of the point $P$ are given by :

$$
\begin{align*}
& X=(r+h) \cos (\varphi) \cos (\lambda)  \tag{6.1-3a}\\
& Y=(r+h) \cos (\varphi) \sin (\lambda)  \tag{6.1-3b}\\
& Z=\left[\left(1-\mathrm{e}^{2}\right) r+h\right] \sin (\varphi) \tag{6.1-3c}
\end{align*}
$$

[^0]
$\varphi$ : geodetic latitude of point $P$
$\lambda$ : geodetic longitude of both P and $\mathrm{P}_{1}$
$h$ : altitude normal to reference oblate spheroid
a : oblate spheroid equatorial radius $=6,378,137 \mathrm{~m}$ based on World Geodetic System 1984 (WGS-84) [43]
e : eccentricity of reference oblate spheroid $\mathbf{e}=0.0818191908426214957($ WGS-84)
b : oblate spheroid polar radius $=\mathbf{a} \sqrt{1-\mathbf{e}^{2}}$
$\theta$ : geocentric latitude of point $P$
$\psi$ : geocentric latitude of point $\mathrm{P}_{1}$

Figure 6.1-1 Oblate Spheroid Earth Model (figure from [24])
where

$$
\begin{equation*}
\mathrm{r}=\frac{\mathrm{a}}{\sqrt{\left(1-\mathrm{e}^{2} \sin ^{2} \phi\right)}} \tag{6.1-3d}
\end{equation*}
$$

$\mathrm{h}=0$ for the beacon transmitter

Given the sub-satellite geocentric latitude and longitude $\left(\theta_{0}, \lambda_{0}\right)$, the satellite ECEF coordinates ( $X_{0}, Y_{0}, Z_{0}$ ) can be found using equations (6.1-2) and (6.1-3), where the altitude above the earth (h) can be approximated to the difference between the satellite orbit radius and the earth semi-major axis $a$. Because the sub-satellite latitude is small, this approximation has little impact on the overall result. The range from point $P$ to the satellite is then given by :

$$
\begin{equation*}
\mathrm{R}=\sqrt{\left(\mathrm{X}_{0}-\mathrm{X}\right)^{2}+\left(\mathrm{Y}_{0}-\mathrm{Y}\right)^{2}+\left(\mathrm{Z}_{0}-\mathrm{Z}\right)^{2}} \tag{6.1-4}
\end{equation*}
$$

So by trial and error, the point on the earth in terms of latitude and longitude which generates a given differential range for two given satellites can be found using equations (6.1-3) and (6.1-4). A similar approach was taken in [7] to find the user's location based on differential ranging from three satellites and using spherical geometry.

A computer program to perform this computation is given in Appendix A with an example of its output. The accuracy of the solution given by this program is limited to approximately $10^{-7}$ degrees ( $3.6 \times 10^{-4}$ seconds) or by the accuracy imposed by the user in terms of allowed differential range error.

### 6.2 GEOSTATIONARY SATELLITE MOTION

### 6.2.1 Impact on Position Accuracy

In practice the geostationary satellites are never perfectly stationary with respect to the rotating earth as discussed in Section 4.1. Even if a satellite could be placed on a synchronous orbit with zero eccentricity, zero inclination and zero longitude drift rate, the satellite motion would soon depart from this geostationary condition because of the orbit perturbations [18]. For instance, the gravitational pull of the sun and moon can change the orbit inclination by $0.005 \%$ day, and the earth's oblateness can give a longitude acceleration of $0.0016 \% / \mathrm{day}^{2}$. In general, the satellites are kept to their nominal position within a given north-south and east-west station keeping errors. As mentioned in Section 3.0, the GOES satellites are maintained at their nominal position to within $\pm 0.1^{\circ}$. This specification does not however take into account the range variation of the satellite. In practice, the satellite motion prediction and analysis is a very complex task. Slabinsky in [18] analyzed a simplified scenario where some perturbations of the sun and moon were neglected. He basically came out with two useful graphs to estimate the variation of the satellite range as a function of orbit eccentricity and inclination. Figures $6.2 .1-1$ and 6.2.1-2 report these graphs. They give the range variation for a satellite to ground station longitude difference $\left(L_{m}-L_{e}\right)$ and ground station latitude. The worst case range variation is given by :

$$
\delta R_{\max }=\left|A_{i}\right|+A_{e}
$$

where $A_{i}$ is the range variation due to the orbit inclination and $A_{e}$ is the one due to the orbit eccentricity. Assuming a worst case inclination of $0.1^{10}$ and an eccentricity of $3.6 \times 10^{-4}$ for the GOES satellite as in Section 4.1, the maximum range variation can be found to be given by :

$$
\begin{aligned}
& \left(\mathrm{A}_{\mathrm{i}}\right)_{\max }=111.3 * 0.1=11.13 \mathrm{~km} \\
& \left(\mathrm{~A}_{\mathrm{e}}\right)_{\max }=43.58 * 0.36=15.69 \mathrm{~km}
\end{aligned}
$$

Having determined the maximum range error when considering the satellites to be perfectly stationary, let us look at the impact on the beacon transmitter location. There are two ways this car be examined. First, a change in the satellite position from the nominal one will result in another position line. So, it is possible to visualize graphically the impact of the satellite motion. However, this will be of little use because the scale used for the map is so large that two position


Figure 6.2.1-1 Amplitudes for inclination part of variations, $\mathrm{nA}_{\mathrm{i}}$ for range-rate and $\mathrm{A}_{\mathrm{i}}$ for range. (Notes: For south latitudes, take negative of quantity for corresponding north latitude. For other inclinations $\mathbf{i}$, multiply quantity from graph by $\mathbf{i}$ degrees.) (from [18])


Figure 6.2.1-2 Amplitude for eccentricity part of variations, $\mathrm{nA}_{e}$ for range-rate and $\mathrm{A}_{\mathrm{e}}$ for range. (Notes: For other eccentricities e, multiply quantity from graph by 1000 e). (from [18])
lines a few kilometers away will be very close to each other. A better approach is to compute the distance between the two lines.

The distance between any two points on the earth has been modeled by several investigators. Ludvik [25] gives two Fortran computer programs to compute the geodetic distance between two points for two techniques i.e. Bowring's inverse algorithm and Vincenty's direct algorithm. Both techniques are accurate, but Vincenty's is slightly better. These programs have been adapted to a Macintosh personal computer (see Appendix B) and have been tested using the known true distance between a set of points as described in [26].

With satellites located at $135^{\circ} \mathrm{W}$ and $75^{\circ} \mathrm{W}$, two position lines were generated which corresponded to 200 km and 215 km range difference. The shortest distance between two points on the same parallel and lying on these position lines was computed. Table 6.2.1-1 shows the results for a few points. These results clearly show that the range difference directly impacts on the accuracy of the position location and that the positioning accuracy is on the same order as the range difference for a satellite spacing of $60^{\circ}$ and a range difference of 200 km (a detailed analysis of the position dilution of precision (PDOP) factor is presented in Section 8.0). It is clear from the above results that the satellite positions must be monitored at all time.

| Latitude <br> (degrees) | Longitude (degrees) |  | Distance (km) |
| :--- | :--- | :---: | :---: |
|  | 0 km range error | 15 km range error |  |
| 22.96 | 103.2756348 | 103.1462402 | 13.27 |
| 45.92 | 102.6359253 | 102.4584656 | 13.77 |
| 72.16 | 99.3142700 | 98.8860779 | 14.65 |

[^1]
### 6.2.2 Satellite Position Tracking

In the previous section, it was shown that the geostationary satellite motion could induce a significant amount of errors on the position location of the beacon and it was concluded that the satellite positions must be monitored. There are several ways to track the geostationary satellite motion. The satellite provider usually keeps track on a monthly basis (if not more frequently) of the satellite position and fires the thrusters to maintain the satellite on its orbit every now and then (corrections are typically required every 10-15 days for a station keeping of $0.1^{\circ}$ in both axis). Although constant communications could be maintained with the satellite provider, it is worthwhile to look at other techniques which could be used.

First of all, we have available the equipment to perform differential ranging and it is intuitively attractive to look at techniques to use this resource to locate the satellites. If reference beacons with exactly known locations are used, then the measured differential range could be used to determine the satellite positions. There are various techniques that can be used to do so.

A straightforward technique is to revert to the process discussed in the previous sections i.e. use known reference beacon locations to find the position of the satellites. Six reference stations are required because we use two satellites at a time and six equations are required to solve for the six unknown variables i.e. the ( $x, y, z$ ) coordinates of the two satellites.

If we denote the satellite coordinates by $\left(\mathrm{X}_{\mathrm{i}}, \mathrm{Y}_{\mathrm{i}}, \mathrm{Z}_{\mathrm{i}}\right), \mathrm{i}=1,2$ and the six reference station coordinates by $\left(\mathrm{x}_{\mathrm{i}}, \mathrm{y}_{\mathrm{i}}, \mathrm{z}_{\mathrm{i}}\right) \mathrm{i}=1, . .6$, then we have :

$$
\Delta R_{i}=\sqrt{\left(X_{2}-x_{i}\right)^{2}+\left(Y_{2}-y_{i}\right)^{2}+\left(Z_{2}-z_{i}\right)^{2}}-\sqrt{\left(X_{1}-x_{i}\right)^{2}+\left(Y_{1}-y_{i}\right)^{2}+\left(Z_{1}-z_{i}\right)^{2}}
$$

for $i=1, \ldots, 6$ where $\Delta R_{i}$ is the measured differential range. This set of nonlinear equations can be solved by using a Taylor series expansion and by considering only the first order terms. If we define:

$$
F_{i}=\Delta R_{i}-\sqrt{\left(X_{2}-x_{i}\right)^{2}+\left(Y_{2}-y_{i}\right)^{2}+\left(Z_{2}-z_{i}\right)^{2}}+\sqrt{\left(X_{1}-x_{i}\right)^{2}+\left(Y_{1}-y_{i}\right)^{2}+\left(Z_{1}-z_{i}\right)^{2}}
$$

then in matrix notation, we have to solve the following set of equations :

$$
\left[\begin{array}{cccccc}
\frac{\partial \mathrm{F}_{1}^{0}}{\partial \mathrm{X}_{1}} & \frac{\partial \mathrm{~F}_{1}^{0}}{\partial \mathrm{Y}_{1}} & \frac{\partial \mathrm{~F}_{1}^{0}}{\partial \mathrm{Z}_{1}} & \frac{\partial \mathrm{~F}_{1}^{0}}{\partial \mathrm{X}_{2}} & \frac{\partial \mathrm{~F}_{1}^{0}}{\partial \mathrm{Y}_{2}} & \frac{\partial \mathrm{~F}_{1}^{0}}{\partial \mathrm{Z}_{2}}  \tag{6.2.2-3}\\
\vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\
\frac{\partial \mathrm{~F}_{6}^{0}}{\partial \mathrm{X}_{1}} & \frac{\partial \mathrm{~F}_{6}^{0}}{\partial \mathrm{Y}_{1}} & \frac{\partial \mathrm{~F}_{6}^{0}}{\partial \mathrm{Z}_{1}} & \frac{\partial \mathrm{~F}_{6}^{0}}{\partial \mathrm{X}_{2}} & \frac{\partial \mathrm{~F}_{6}^{0}}{\partial \mathrm{Y}_{2}} & \frac{\partial \mathrm{~F}_{6}^{0}}{\partial \mathrm{Z}_{2}}
\end{array}\right]\left[\begin{array}{c}
\mathrm{dX}_{1} \\
\mathrm{dY}_{1} \\
\mathrm{dZ}_{1} \\
\mathrm{dX}_{2} \\
\mathrm{dY}_{2} \\
\mathrm{dZ}_{2}
\end{array}\right]=\left[\begin{array}{c}
\mathrm{F}_{1} \\
\mathrm{~F}_{2} \\
\mathrm{~F}_{3} \\
\mathrm{~F}_{4} \\
\mathrm{~F}_{5} \\
\mathrm{~F}_{6}
\end{array}\right]
$$

where $\partial \mathrm{F}_{\mathrm{i}}{ }^{0} \partial \mathrm{U}_{\mathrm{j}}$ is $\partial \mathrm{F}_{\mathrm{i}} / \partial \mathrm{U}_{\mathrm{j}}$ evaluated at $\left(\mathrm{X}_{\mathrm{k}}{ }^{0}, \mathrm{Y}_{\mathrm{k}}{ }^{0}, \mathrm{Z}_{\mathrm{k}}{ }^{0}\right), \mathrm{k}=1,2$ which is an initial good estimate of the satellite position (e.g. their nominal positions). Solving this equation gives ( $\mathrm{dX}_{\mathrm{k}}, \mathrm{dY}_{\mathrm{k}}, \mathrm{d} \mathrm{Z}_{\mathrm{k}}$ ) $\mathrm{k}=1,2$ which can be used to form the next best estimate i.e. :

$$
\begin{aligned}
& X_{k}{ }^{1}=X_{k} 0+d X_{k} \\
& Y_{k^{1}}=Y_{k} 0+d Y_{k} \\
& Z_{k}{ }^{1}=Z_{k} 0+d Z_{k} \quad \mathrm{k}=1,2
\end{aligned}
$$

The process can then be repeated with each new estimate. Because the solution is convergent, the iterations can be stopped when the dX 's, dY 's and dZ's are less than $\varepsilon$. Based on the computer results of several examples, it has been noticed that only a few repeats are required (e.g. typically 3-5).

When random differential range measurement errors are introduced, the set of equations can still provide a solution but the solution only represents the satellite positions with the etroneous differential ranges. The amount of errors in the satellite position is a function of the differential range error and the geometry made by the satellites and the reference stations. The latter is referred to as the position dilution of precision (PDOP) factor which will be developed later on for the beacons. Briefly, it is the factor that must multiply the range error to obtain the user position error. This factor is completely determined by the geometry of the system.

The analytical equations for determining the PDOP as a function of the satellite and reference station locations have not been derived. However, it is easy to understand conceptually that the PDOP factor will be very large due to the fact that the satellites view the reference stations ${ }^{l}$ lying in $^{\text {in }}$ a very small cone with very little spatial discrimination. Accordingly, even if the reference
stations are selected to give the best PDOP, computer simulations of a few cases have showed that the PDOP is between 10-80. So a differential range measurement of 100 meters (about the minimum in the differential mode) would result in a satellite position error of approximately 18 km ! Clearly this is not satisfactory and other techniques must be considered.

One possible alternate technique is to use an absolute range measurement instead of a differential one. This requires the reference stations and a master station to be time synchronized which is difficult to achieve but this problem was left aside temporarily. Unfortunately, again it has been found that the large PDOP was the main detrimental element which kept the accuracy above 1 km for a 100 meters range error. Other techniques were also investigated but none provided any better solution than the 1000 meters error on the satellite positions. Such an error is intolerable and we must rely on the satellite provider to make available the satellite position information with enough accuracy.

The satellite provider is equipped with sophisticated tools such as frequent bearing and range measurements, orbit models, Kalman filters, etc. to keep track of the satellite position. For our application, it would be ideal to receive continuously from the satellite (e.g on a beacon signal) its current estimated position. Because the satellite provider has all the tools to compute and predict the satellite position with very high accuracy, it will require a minimum effort on their part to make this information available to all users. As a minimum, the availability of the ephemeris data can be useful for our application but it would still require a significant amount of processing to come out with good predicted satellite positions.

The accuracy to which one can estimate the satellite position depends strongly on how often measurements are made and how sophisticated is the orbit model. In [7] and [21] the satellite position is assumed to be predicted to within 20 m in all axis. Although it is not explicitly explained how this number is derived, we will assume in the following that this accuracy can be achieved.

### 6.3 CHANNEL DELAYS

The terminal measures the time delay between the signals received from different satellites and to find the range difference, this time delay is divided by the wave velocity in the channel. The wave velocity in the atmosphere is often approximated to $2.9979 \times 10^{8} \mathrm{~m} / \mathrm{s}$ but in practice it is not constant with time. There are two layers of the atmosphere which contribute to this velocity variation. They are the ionosphere and the troposphere.

## 63,1 The Ionosphere Delay

The range error assuming a wave velocity c for the ionosphere is given by [41]:

$$
\begin{equation*}
\Delta \mathrm{R}=\frac{40.3}{\mathrm{f}^{2}} \mathrm{TEC} \tag{6.3.1-1}
\end{equation*}
$$

Where TEC is the total electron content along the path in electrons $/ \mathrm{m}^{2}$, and f is the frequency of interest in Hertz.

Alternately equation (6.3.1-2) can be rewritten as :

$$
\begin{equation*}
\Delta \mathrm{R}=\frac{40.3}{\mathrm{f}^{2}} \mathrm{f}(\theta) \mathrm{TEC}_{\mathrm{v}} \tag{6.3.1-2}
\end{equation*}
$$

where TEC $_{\mathrm{v}}$ is the vertical total electron content and :

$$
\begin{equation*}
f(\theta)=\sec \left[\sin ^{-1}\{0.94792 \cos (\theta)\}\right] \tag{6.3.1-3}
\end{equation*}
$$

and $\theta$ is the elevation angle. The factor $f(\theta)$ varies between 1 at zenith and 3.1 at $0^{\circ}$ elevation angle.

The true range is always less than the one measured assuming a velocity ${ }^{c}=2.9979 \times 10^{8} \mathrm{~m} / \mathrm{s}$ and accordingly $\Delta \mathrm{R}$ is always positive. The TEC is a function of many variables including short and long term changes in solar ionizing flux, magnetic activity, season, time of day, user location and viewing angle. The TEC varies typically between $10^{16}$ and ${ }^{1019} \mathrm{el} / \mathrm{m}^{2}$. Figure 6.3.1-1 shows the time delay introduced by the ionosphere for various values of TEC. At 406 MHz , the range error is between 2.44 m and 2.44 km . However it is mentioned


Figure 6.3.1-1 Ionospheric Time Delay Versus Frequency for Various Values of Electron Content (from [27])
in [28] that a typical TEC value is $2 \times 10^{17} \mathrm{el} / \mathrm{m}^{2}$ corresponding to a delay of 49 m . In addition, the TEC exceeds $5 \times 10^{17} \mathrm{el} / \mathrm{m}^{2}$ approximately $10 \%$ of the time and it rarely exceeds $10^{18} \mathrm{el} / \mathrm{m}^{2}$ which clearly limits the delay to 244 m for all practical cases in the mid to upper latitude regions.

The TEC spatial variation is relatively smooth. Thus if the two satellites are close to each Other, the beacon signal for each satellite can be assumed to go through the same ionospheric region with relatively constant TEC. Because we are interested in the time difference of arrival, the ionospheric delay would then cancel out. Table 6.3.1-1 from [29] shows the amount of ionospheric delay that does not cancel out as a function of the station separation when the stations are looking at the same GPS satellite. Our application is slightly different but these results can be used to estimate the amount of delay cancellation using the differential technique.

Using the relationships given in equations (6.3.1-2) and (6.3.1-3), we see that the residual delay after differential timing is given by :

$$
r=\frac{\operatorname{TEC}_{v 2} f\left(\theta_{2}\right)}{\operatorname{TEC}_{v 1} f\left(\theta_{1}\right)}=\frac{g(d) \operatorname{TEC}_{v 1} f\left(\theta_{2}\right)}{\operatorname{TEC}_{v 1} f\left(\theta_{1}\right)}=g(d) \frac{f\left(\theta_{2}\right)}{f\left(\theta_{1}\right)}
$$

Where $\mathrm{g}(\mathrm{d})$ is a factor depending on the distance (d) between the two stations. Whether $\mathrm{g}(\mathrm{d})$ is a linear function of distance or not, we can postulate that if the distance between the two signals passing through the ionosphere remains constant, then $g(d)$ is constant and the residual delay relative to the GPS case is increased by the ratio of $f\left(\theta_{2}\right) / f\left(\theta_{1}\right)$. For instance, let us consider a simple case where the beacon is located on the equator exactly between the two satellites spaced by $\alpha$ degrees, then their elevation angles are given by

$$
\theta_{\mathrm{i}}=\tan ^{-1}\left[\frac{\cos (\alpha / 2)-\frac{\mathrm{R}_{\mathrm{e}}}{\mathrm{R}_{\mathrm{s}}}}{\sin (\alpha / 2)}\right] \quad \mathrm{i}=1,2
$$

Where $\mathrm{R}_{\mathrm{e}}$ and $\mathrm{R}_{\mathrm{S}}$ are defined in Section 5.0.
When the beacon is moved along the equator, it can be shown that the difference between the elevation angles will increase up to a maximum when one of the satellite is at the horizon. For a ${ }^{300}$ spacing between satellites, this maximum is approximately $39^{\circ}$. When the beacon is moved ${ }^{\text {along the }} 45^{\text {th }}$ parallel, this maximum is reduced to approximately $20^{\circ}$. So, for location techniques of interest to Canada and for a given distance between the signals passing through the ionosphere,

| Station Separation |  | Residual Delay <br> $(\%)$ |
| :---: | :---: | :---: |
| $(\mathrm{nmi})$ | $(\mathrm{km})$ |  |
| 0 | 0 | 0 |
| 1 | 1.852 | 2 |
| 10 | 18.52 | 8 |
| 50 | 92.60 | 17 |
| 100 | 185.2 | 24 |
| 500 | 926.0 | 52 |
| 1000 | 1852.0 | 71 |
| 2000 | 3704.0 | 91 |

Table 6.3.1-1 Differential Ionospheric Delay Reduction (from [29])

| Distance Between Signals Over <br> the Ionosphere |  |
| :---: | :---: |
| $(\mathrm{nmi})$ | Residual Delay <br> $(\%)$ |
| 0 | 0 |
| 1 | 1.852 |
| 10 | 18.52 |
| 50 | 92.60 |
| 100 | 185.2 |
| 500 | 926.0 |
| 1000 | 3704.0 |

Table 6.3.1-2 Minimum Differential Ionospheric Delay Reduction for the Time Difference of Arrival Location Technique Assuming a $30^{\circ}$ Satellite Spacing and for Canada
the residual delay given in Table 6.3.1-1 should be increased by a factor of $f\left(20^{\circ}\right)=2.2$ and up to a maximum residual delay of $100 \%$ meaning total uncorrelation. Table 6.3.1-2 shows the modified values.

The distance between the two signal paths passing through the ionosphere is a function of the elevation angle and satellite spacing. This distance is minimum for a beacon located on the equator. Referring to Figure 6.3.1-2 which represents the simple case when the beacon is located midway between the two satellites and on the equator, we find that the distance $d$ is given by :

$$
\begin{aligned}
& d=2 \beta\left(\mathrm{R}_{\mathrm{e}}+\mathrm{R}_{\text {iono }}\right) \\
& \beta=2 \tan ^{-1}\left[\frac{\tan \left(\frac{a+b}{2}\right) \cos \left(\frac{A+B}{2}\right)}{\cos \left(\frac{A-B}{2}\right)}\right] \\
& b=\sin ^{-1}\left(\frac{\sin a}{\sin A}\right) \\
& a=R_{\text {iono }} / R_{e} \\
& B=90^{\circ}
\end{aligned}
$$

Where $R_{\text {iono }}$ is the ionosphere mean altitude above the earth (assumed to be 350 km ), $\mathrm{R}_{\mathrm{e}}$ is the earth radius and $A, B$ are the angles opposite to sides $a$ and $b$ respectively defined in Figure 6.3.1-2. Note that the above trigonometric functions must be performed in radians. With a satellite spacing of $30^{\circ}$ we get $A=72.375^{\circ}, a=0.05494, b=0.057617, d=217.24 \mathrm{~km}$. Table 6.3.1-2 shows that between 52 and $100 \%$ of the delay is not cancelled for a distance of 217 km . This represents the best case for Canada where the distance increases with the latitudes, and it is clear that a model to estimate the ionospheric delay is required.

It is possible to estimate the TEC or to eliminate it in equation (6.3.1-1) using a second frequen frequency. Indeed, when we apply equation (6.3.1-1) to two signals with different frequencies, we have :


Figure 6.3.1-2 Model for the Discussion of the Ionospheric Delay Spatial Correlation

$$
\begin{align*}
& \Delta \mathrm{t}_{2}=\frac{40 \cdot 3}{\mathrm{cf}} \mathrm{2} \\
& \Delta \mathrm{t}_{1}=\frac{40.3}{\mathrm{Cf}} \mathrm{~T}  \tag{6.3.1-4}\\
& \mathrm{TEC} \\
& \delta \mathrm{t}=\Delta \mathrm{t}_{2}-\Delta \mathrm{t}_{1}
\end{align*}
$$

Where $\Delta t_{1}$ is the time delay at frequency $f_{1}$ and $\delta t$ is the time delay between the two signal frequencies transmitted by a single receiver at the same time and measured at a receiver. Rearranging the above equations we find that :

$$
\begin{equation*}
\Delta t_{1}=\frac{f_{2}^{2}}{f_{1}^{2}-f_{2}^{2}} \delta t \tag{6.3.1-5}
\end{equation*}
$$

Knowing the time delay introduced by the ionosphere ( $\Delta t_{1}$ ), we can find the range error $\Delta R$ introduced by the ionosphere and make the appropriate correction. A second frequency is however not available in the search and rescue system.

When only a single frequency signal is available, one is required to predict, as much as possible, the range error based on ionospheric conditions. A great deal of effort has been deployed to find such models for the Global Positioning System (GPS) operating at 1.5 GHz . Klobuchar [30] mentioned the use of a relatively simple model where approximately $50 \%$ of the error introduced by the ionosphere can be corrected. With a state of the art model, up to 70 to $80 \%$ can be compensated. Given that the time delay varies approximately with $1 / \mathrm{f}^{2}$ for $\mathrm{f}>100 \mathrm{MHz}$, it is possible to use the GPS single frequency correction technique discussed in [30]. This technique however requires the use of some coefficients transmitted by the GPS satellites. These coefficients Which required updating about every ten days are used to model the amplitude of the TEC variation. It would therefore be possible to read these coefficients from the GPS signal and apply the correction factor determined by the technique discussed in [30].

From the above discussion, it is clear that the best approach is to rely on the GPS technique and to scale the results by a factor of $(1575.42 / 406)^{2}=15.057$. Although this approach means that ${ }^{\text {a }}$ GPS receiver must be available, it is considered significantly simpler than trying to re-derive the parameters in the broadcast channel. The GPS system can provide an ionospheric rms range error of $4.5-10 \mathrm{~m}$ [31] which translates at 406 MHz to a maximum rms range error of $\sigma_{\text {iono }}=150 \mathrm{~m}$. $N_{\text {ote }}$ again that in the absence of an ionospheric model, we could expect an error of 244 m for
most of the time. These errors will be used in Section 9.0 to derive the overall positioning accuracy.

### 6.3.2 The Troposphere Delay

The troposphere range error is caused by two effects : angular bending of the waves which increases the path length with reference to free space and a decrease in propagation velocity. Both effects result from a change in the refraction index as a function of altitude in the troposphere and are essentially independent of frequency up to 30 GHz . The range correction factor is fully described in [32] and after some simplifications we obtained :

$$
\begin{equation*}
\Delta \mathrm{R}=\mathrm{f}(\theta) * \Delta \mathrm{R}(\mathrm{~h}) \tag{6.3.2-1}
\end{equation*}
$$

where $\Delta R(h)$ is given by :
for $0<\mathrm{h} \leq 1 \mathrm{~km}$

$$
\begin{equation*}
\Delta R(h)=\left(2464.4042-324.8 \mathrm{~h}-22.39578 \mathrm{~h}^{2}\right) \quad \mathrm{mm} \tag{6.3.2-1a}
\end{equation*}
$$

for $1<\mathrm{h} \leq 9 \mathrm{~km}$

$$
\begin{equation*}
\Delta R(h)=\left[2283.7805 \exp \left(\frac{1-\mathrm{h}}{8.1561}\right)-124.3926\right] \mathrm{mm} \tag{6.3.2-1b}
\end{equation*}
$$

for $9 \mathrm{~km}<\mathrm{h} \leq \mathrm{h}_{\text {sat }}$

$$
\begin{equation*}
\Delta \mathrm{R}(\mathrm{~h})=[2656.26 \exp (-0.1424 \mathrm{~h})] \mathrm{mm} \tag{6.3.2-1c}
\end{equation*}
$$

and $f(\theta)$ is given by :

$$
\begin{equation*}
\mathrm{f}(\theta)=\frac{1}{\sin \theta+\frac{0.00143}{\tan \theta+0.0455}}=\frac{1}{\sin \theta} \tag{6.3.2-2}
\end{equation*}
$$

where $h$ is the altitude of the transmitter (in kilometers) above the sea level, $\theta$ is the elevation angle to the satellite and $\mathrm{h}_{\text {sat }}$ is the satellite altitude. Because the altitude above the sea level does not
exceed approximately 4 km in Canada, only equations (6.3.2-1a) and (6.3.2-1b) are of interest here. In general, h is not known a priori, however, if required it can be estimated using an electronic data base. The position is then estimated using $\mathrm{h}=0$, the altitude for that estimated position is found from the data base and calculations are then repeated with a new $h$. In the following, we assume $h=0$ and neglect the error introduced which is about 1 m for the worst case in Canada.

For the tropospheric range error, a residual error of less than 4 m can be expected for $95 \%$ of the time [31] with the above model and exact h . A bias of 1 m can be assumed for the tropospheric model when $h$ is assumed zero for Canada. Looking at the above equation, it can be found that an error as large as 78 meters ( $\theta=0$ and $\mathrm{h}=0$ ) can be expected if the troposphere model is not implemented.

## Z.0 ESTIMATION OF THE TIME DIFFERENCE OF ARRIVAL

Up to now, we have looked at all the external perturbations of time difference of arrival estimates. It is time to look at the accuracy of the time of arrival estimation itself. The problem consists of determining the delay between a single signal relayed via two satellites to a control ground station. The content of the two received signals is identical but the phase and amplitude may be different. In addition, there may be other signals present with signal characteristics close to the ones of interest. In the following, we first determine the maximum time difference that can be expected. Then, we discuss the problem of time difference estimation relayed via two satellites.

It is assumed that the control ground station receiver for each satellite link is as illustrated in Figure 7.0-1 where the frequency down-converter is locked to the satellite beacon and therefore the frequency offset between two or more satellites can be assumed to be zero. The signal at the input of the delay estimator is assumed centred on a 5 kHz IF which corresponds to the minimum intermediate frequency as required by the long term frequency stability of the beacon transmitter.

### 7.1 MAXIMUM TIME DIFFERENCE OF ARRIVAL

The maximum time difference of arrival is a function of the satellite spacing. In order to get an idea of the magnitude of the maximum delay, we use the spherical geometry introduced in Section 5.0. The maximum delay between a signal received via two satellites occurs for a station When one of the satellites is seen with the minimum operating elevation angle. The elevation angle when the earth is assumed a sphere is given by :

$$
\begin{equation*}
\theta=\tan ^{-1}\left[\frac{\cos \beta-R_{c} / R_{s}}{\sin \beta}\right] \tag{7.1-1}
\end{equation*}
$$

Where all variables have been defined in Section 5.0 under equations (5.0-1a), (5.0-1b) and (5.0${ }^{l}$ c). Using the following trigonometric identity :

$$
\begin{equation*}
\sin \left(\tan ^{-1} \mathrm{x}\right)=\frac{\mathrm{x}}{\sqrt{1+\mathrm{x}^{2}}} \tag{7.1-2}
\end{equation*}
$$



Figure 7.0-1 Basic Satellite Receiver Configuration
and applied to equation (5.0-1a) gives for the range to each satellite :

$$
\begin{equation*}
R_{i}=R_{e} \sqrt{\left(\frac{R_{s}}{R_{e}}\right)^{2}-2 \frac{R_{s}}{R_{e}} \cos \beta_{i}+1} \quad i=1,2 \tag{7.1-3}
\end{equation*}
$$

where $\cos \left(\beta_{i}\right)=\cos (\varphi) \cos \left(\lambda-\lambda_{i}\right)$
$\lambda_{i}$ is the longitude of satellite \#i
( $\varphi, \lambda_{\mathrm{i}}$ ) are the coordinates of the beacon transmitter.
The differential range is then given by :

$$
\begin{equation*}
\Delta R=R_{1}-R_{2} \tag{7.1-4}
\end{equation*}
$$

Where the satellite \#1 is assumed to be seen by the beacon transmitter with an elevation angle $\theta_{1}=$ $\theta_{\min }$. For a given $\theta_{\min }$ we can solve equation (7.1-1) to obtain the corresponding $\beta$. For $\theta_{1}=$
$\theta_{\min }=50$ $\theta_{\min }=5^{\circ}$ we get $\beta_{1}=76.33855^{\circ}$ which when applied to equations (7.1-3) and (7.1-4) gives :

$$
\begin{align*}
R_{1} & =6.4544 R_{e} \\
\Delta R & =R_{e}\left[6.4544-\sqrt{\left(\frac{R_{s}}{R_{e}}\right)^{2}-2 \frac{R_{s}}{R_{e}} \cos \beta_{2}+1}\right] \tag{7.1-5}
\end{align*}
$$

This last equation is maximized when $\cos \left(\beta_{2}\right)$ is maximized. Using equation (5.0-1b) applied to both satellites, we have :

$$
\cos \left(\beta_{1}\right)=\cos (\varphi) \cos \left(\lambda-\lambda_{1}\right) \quad \text { from which } \quad \cos (\varphi)=\frac{\cos \left(\beta_{1}\right)}{\cos \left(\lambda-\lambda_{1}\right)}
$$

and

$$
\begin{aligned}
\cos \left(\beta_{2}\right) & =\cos (\varphi) \cos \left(\lambda-\lambda_{2}\right)=\frac{\cos \left(\beta_{1}\right)}{\cos \left(\lambda-\lambda_{1}\right)} \cos \left(\lambda-\lambda_{2}\right) \\
& =\frac{\cos \left(\beta_{1}\right)}{\cos \left(\lambda-\lambda_{1}\right)}\left\{\cos \lambda \cos \lambda_{2}+\sin \lambda \sin \lambda_{2}\right\} \\
& =\frac{\cos \left(\beta_{1}\right)}{\cos \left(\lambda-\lambda_{1}\right)}\left\{\left[\cos ^{2} \lambda_{1}+\sin ^{2} \lambda_{1}\right]\left(\cos \lambda \cos \lambda_{2}+\sin \lambda \sin \lambda_{2}\right)\right\}
\end{aligned}
$$

using basic trigonometric identities. After some manipulations we get :

$$
\begin{equation*}
\cos \beta_{2}=\left(\cos \beta_{1}\right)\left[\cos \left(\lambda_{1}-\lambda_{2}\right)-\tan \left(\lambda-\lambda_{1}\right) \sin \left(\lambda_{1}-\lambda_{2}\right)\right] \tag{7.1-6}
\end{equation*}
$$

The above equation is maximized when $\left|\left(\lambda-\lambda_{1}\right)\right|$ is maximized and :

$$
\begin{aligned}
& \left(\lambda_{1}-\lambda_{2}<0 \text { and } \lambda-\lambda_{1}>0\right) \text { or }\left(\lambda_{1}-\lambda_{2}>0 \text { and } \lambda-\lambda_{1}<0\right) \\
& \text { With } \cos \left(\beta_{1}\right)=\cos (\varphi) \cos \left(\lambda-\lambda_{1}\right)
\end{aligned}
$$

the maximization of $\left|\left(\lambda-\lambda_{1}\right)\right|$ is equivalent to the minimization of $\cos \left(\lambda-\lambda_{1}\right)$ for a given $\beta_{1}$. This is achieved when $\cos (\varphi)$ is maximum i.e. at $\varphi=0$ when the beacon is located on the equator. In this case, $\left(\lambda-\lambda_{1}\right)_{\max }=\beta_{1}$ and equation (7.1-6) becomes :

$$
\begin{equation*}
\left.\left(\cos \beta_{2}\right)_{\max }=\left(\cos \beta_{1}\right)\left[\cos \left(\lambda_{1}-\lambda_{2}\right)+\tan \beta_{1} \sin \left(\lambda_{1}-\lambda_{2}\right)\right)\right] \tag{7.1-7}
\end{equation*}
$$

Substituting equation (7.1-7) into equation (7.1-5), we can find the maximum range difference as a function of the satellite spacing $\left|\lambda_{1}-\lambda_{2}\right|$. Table $7.1-1$ shows the maximum time difference of arrival as a function of satellite spacing. Because Canada spans approximately $140^{\circ}$ in longitude, it is fair to assume that the spacing between two adjacent satellites will not exceed $70^{\circ}$ which corresponds to a maximum delay of about 18 ms .

| Satellite Spacing <br> (degrees) | Max. Delay <br> $(\mathrm{ms})$ |
| :---: | :---: |
| 5 | 1.84 |
| 10 | 3.65 |
| 15 | 5.41 |
| 20 | 7.11 |
| 25 | 8.74 |
| 30 | 10.28 |
| 35 | 11.71 |
| 40 | 13.02 |
| 45 | 14.20 |
| 50 | 15.22 |
| 55 | 16.09 |
| 60 | 16.79 |
| 65 | 17.31 |
| 70 | 17.64 |
| 75 | 17.79 |

Table 7.1-1 Maximum Delay as a Function of Satellite Spacing When the Minimum Operating Elevation Angle is Set at $5^{\circ}$.

### 7.2 DELAY ESTIMATION

There are basically two tasks that must be performed in the time difference of arrival processor. First, we have to detect that a beacon signal is present and second we have to estimate the delay between the three received signals. The detection problem has been addressed in a previous project where a detector based on a spectral estimator has been designed and has been shown to perform very well in an operational scenario (see [33]). Although we could address the problem in a general sense i.e. treat the problem as a detection and estimation problem, it is in general, less difficult to design an optimum processor for either detection or estimation than to design the combined one. The loss of optimality is in general low for such an approach. So in the following it is assumed that the detection process has been performed and we have been given an indication that a signal has been detected. In addition, we will further assume that three detectors are available to process the signal received from each satellite in parallel. The latter assumption allows us to assume that we know which signal is received first, second and third so that the problem is simplified to an estimation of the delay between signals and this eliminates the uncertainty present when both delay and advance can occurred.

It is important in designing a processor to consider what is known. Although in the ideal situation we would like to have a processor which needs to know very little about the signal itself (this is required for the DND's applications where the interfering signal is not known a priori), the complexity of the processor is generally proportional to the amount of unknowns. For the case of distress beacons localization:
a) we know that the signal starts with a 160 ms un-modulated carrier followed by a known and constant 24 bit pattern;
b) we know the signal period and the statistical law that governs the repetition period;
c) we know that in an operational scenario there could be other signals interfering with the one of interest;
d) we have a "rough" estimate of the frequency of the carrier as derived from the detection processor;
e) based on the knowledge of the frequency, we could refine the estimate of the carrier frequency to 1 ) bring the signal to baseband and reduce the noise bandwidth thus
improving the processor input SNR , and 2) to demodulate the signal to recover the modulation bits;
f) we know that the delay is less than 18 ms ;
g) finally, we know which signal is in advance relative to the other one.

The first five characteristics have been exploited in the GOES processor [33] to perform distress detection. Thus if a processor equivalent to the GOES one is assumed available, we have some means to recover the information bits in addition to the characteristics of $f$ ) and $g$ ).

From the parameter estimation theory, the optimum delay estimator in the Maximum Likelihood (ML) sense for signals embedded in additive white Gaussian noise (AWGN) is given by $\hat{\mathrm{T}}_{\text {which satisfies the following [44] : }}$

$$
\begin{equation*}
\left[\frac{2}{\mathrm{~N}_{\mathrm{o}}} \int_{0}^{\mathrm{T}_{\mathrm{b}}}\left\{\mathrm{r}_{\mathrm{i}}(\mathrm{t})-\mathrm{s}(\mathrm{t}, \mathrm{~T})\right\} \frac{\partial \mathrm{s}(\mathrm{t}, \mathrm{~T})}{\partial \mathrm{T}} \mathrm{dt}\right]_{\mathrm{T}=\hat{\mathrm{T}}}=0 \tag{7.2-1}
\end{equation*}
$$

Where $r_{i}(t)$ is the received signal for satellite $\# i, s(t, T)$ is the noise free signal with delay $T, T_{b}$ is the observation period and $N_{o}$ is the noise power spectral density. Although this equation does not give too much insight on how to implement the estimator, after some manipulations and approximations, it can be shown that this is equivalent to choosing $\widehat{T}$ such that the correlation between $r_{i}(t)$ and $s(t, \widehat{T})$ is maximum.

When applied to the system of interest, the correlation must be performed between the two noisy received signals assuming that we do not know the information in the beacon signal. Figure 7.2-1 shows how such a processor can be implemented digitally. It is clear that the estimator accuracy is limited by the number N of correlators used to cover the 18 ms uncertainty
${ }^{\text {over }}$ a certain ${ }^{\circ}$ ver a certain range of SNR and by noise thereafter. Because the signal SNR of interest is low


Figure 7.2-1 Block Diagram of Optimum Estimator
(e.g. as low as -10 dB for and input filter bandwidth of 10 kHz ), our estimate accuracy will likely be limited by noise.

The above processor is difficult to analyze because we have the product of two noisy terms at the output of the correlator. This problem is similar to the design of an optimum differential demodulator where the approximation used to result in manageable noise statistics consists of assuming that the SNR is large enough to make the product of the noise terms to be negligible relative to the other terms [37]. When making this assumption, the correlator outputs are assumed to contain twice as much noise as in the case of one noisy signal.

The signal-to-noise ratio (SNR) at the input of the correlator is given by :

$$
(S N R)_{\text {in }}=\frac{C}{\text { NoB }_{\text {in }}}
$$

Where $\mathrm{B}_{\mathrm{in}}$ is the input noise bandwidth. The variance of the noise being the noise power, we have:

$$
\sigma_{i}^{2}=\mathrm{NoB}_{\text {in }}=\frac{1}{(\mathrm{SNR})_{\mathrm{in}}}
$$

Where $\mathrm{C}=1$ has been assumed. The correlator can be seen as a filter which reduces the noise bandwidth by the product $\mathrm{B}_{\mathrm{in}} \mathrm{T}_{\mathrm{b}}$ where $\mathrm{T}_{\mathrm{b}}$ is the correlation period. At the output of the correlator, We have a variance equal to :

$$
\left(\sigma_{1}\right)^{2}=\frac{\sigma_{i}^{2}}{\left(\mathrm{~B}_{\mathrm{in}} \mathrm{~T}_{\mathrm{b}}\right)}=\frac{1}{(\mathrm{C} / \mathrm{No}) \mathrm{T}_{\mathrm{b}}}
$$

and when the doubling effect is taken into account, the correlator outputs have a variance given by :

$$
\sigma^{2}=\frac{2}{(\mathrm{C} / \mathrm{No}) \mathrm{T}_{\mathrm{b}}}
$$

Which is independent of the delay. Defining $R_{i}$ as the autocorrelation level for a delay of $i \tau$ and assuming for now that both signals are aligned (i.e. they have no delay), the density function of the ${ }^{\text {correlator outputs ( } \mathcal{F} \text { 's) is given by : }}$

$$
f\left(\mathcal{G}_{\mathrm{i}}\right)=\frac{1}{\sqrt{2 \pi \sigma}} \exp \left\{\frac{-\left(\mathfrak{G}_{\mathrm{i}}-\mathrm{R}_{\mathrm{i}}\right)^{2}}{2 \sigma^{2}}\right\}
$$

The probability that $\mathcal{F}_{\mathcal{i}}$ is greater than $\mathcal{F}_{j}$ for all $j \neq i$ is equivalent to the probability that $z_{\mathrm{ij}}=\mathcal{F}_{\mathrm{i}}-\mathcal{F}_{\mathrm{j}}$ is greater than zero for all $\mathrm{j} \neq \mathrm{i}$. Both $\mathcal{F}_{\mathrm{i}}$ and $\mathcal{F}_{\mathrm{j}}$ being Gaussian, $\mathrm{z}_{\mathrm{ij}}$ is Gaussian with mean $\left(\mathrm{R}_{\mathrm{i}}-\mathrm{R}_{\mathrm{j}}\right)$ and variance $\sigma^{2}\left\{1-\mathrm{R}_{\mathrm{i}-\mathrm{j}}\right\}$. So the probability that $\mathrm{z}_{\mathrm{ij}}$ is greater than zero is given by :

$$
\begin{equation*}
\mathrm{Z}_{\mathrm{ij}}=\operatorname{Prob}\left(\mathrm{z}_{\mathrm{ij}}>0\right)=\mathrm{Q}\left\{\frac{-\left(\mathrm{R}_{\mathrm{i}}-\mathrm{R}_{\mathrm{j}}\right)}{\sigma \sqrt{1-\mathrm{R}_{\mathrm{i}-\mathrm{j}}}}\right\} \tag{7.2-2}
\end{equation*}
$$

where $\mathrm{Q}(\mathrm{x})=\frac{1}{\sqrt{2 \pi}} \int_{\mathrm{x}}^{\infty} \mathrm{e}^{-\mathrm{y}^{2} / 2} \mathrm{dy}$

In the above equations it was assumed that the signals were aligned i.e.there was no delay. When the delay is non-zero, say $k \tau$, the autocorrelation level indices in equation (7.2-2) are replaced by their absolute difference relative to k i.e. :

$$
\begin{equation*}
\mathrm{Z}_{\mathrm{ij} / \mathrm{k}} \equiv \operatorname{Prob}\left(\mathrm{z}_{\mathrm{ij}}>0 / \text { delay is } \mathrm{k} \tau\right)=\mathrm{Z}_{|\mathrm{k}-\mathrm{i}||\mathrm{k}-| |} \tag{7.2-3}
\end{equation*}
$$

The probability of the estimate being within $m \tau$ is then approximated by :

$$
\begin{equation*}
\operatorname{Prob} \text { (error } \leq m \tau) \approx \sum_{i=k-m}^{k+m} \prod_{\substack{j=0 \\ j \neq i}}^{N-1} \operatorname{Prob}\left(z_{i j}>0 / \text { delay is } k \tau\right) \tag{7.2-4}
\end{equation*}
$$

A computer program to solve equations (7.2-2) and (7.2-4) has been developed (see Appendix C). The autocorrelation level used for the integration was given by a typical beacon identification code and some variation may be obtained for different codes. These differences are however expected to be small because we are only interested in the area close to the main peak i.e. when the two signals are aligned and in this area most beacon identification codes exhibit the same autocorrelation pattern.

Figure 7.2-2 shows the results for the case when the spacing between each correlator is either $12 \mu \mathrm{~s}$ or $24 \mu \mathrm{~s}$. The figure shows the probability that the estimator will give no error in the estimated delay or an error of one correlator spacing $(\tau)$ assuming that the real delay is a multiple of $\tau$. From the figure, it is clear that an accuracy better than $12 \mu \mathrm{~s}$ is unlikely for the minimum C/No of $30 \mathrm{~dB}-\mathrm{Hz}$. We can also conclude from the observation of the two sets of curves that sampling faster i.e. taking the correlator spacing smaller than $12 \mu \mathrm{~s}$ will not bring additional confidence in the estimate.

These results have been verified through a Monte Carlo simulation of the estimator (computer program given in Appendix D). For the simulation, it has been assumed that the signal was down-converted to baseband, filtered with a 800 Hz lowpass filter, and sampled at 41.6 kHz $(\tau=24 \mu \mathrm{~s})$ or $83.2 \mathrm{kHz}(\tau=12 \mu \mathrm{~s})$. Then the correlation over 458 ms was computed and the maximum output was selected as representing the estimated delay. Results are shown in Figures $7.2-3$ and $7.2-4$ for $\tau=24 \mu \mathrm{~s}$ and $\tau=12 \mu \mathrm{~s}$ respectively along with the theoretical results. In general the theory and the simulation results agree quite well for $\mathrm{C} / \mathrm{No}$ in excess of $30 \mathrm{~dB}-\mathrm{Hz}$.

Up to now, we considered the case when the delay was a multiple of the correlator spacing. This assumption made the estimate error a multiple of the correlator spacing. In practice the delay could lie anywhere within the correlator spacing such that the previous results indicating No Error should be interpreted as $\pm \tau$ and those indicated as being in error by $\tau$ should in fact be $\pm 2 \tau$ with uniform statistical distribution. In this case, the variance of the estimate is given by :

$$
\left(\sigma_{t}\right)^{2}=\frac{\tau^{2}}{3}
$$

$\mathrm{Al}_{\mathrm{so}}$, the confidence level is indicated in Figure $7.2-2$ for $\tau=24$ or $12 \mu \mathrm{~s}$. For $\tau=12 \mu \mathrm{~s}$, we have $\sigma_{l=6.9} \mu_{\mathrm{s}}$ with a confidence level of approximately $90 \%$ at a C/No of $30 \mathrm{~dB}-\mathrm{Hz}$.

Because the expected delay is bounded by 18 ms , the number of correlation for ${ }^{c}$ mputation is therefore $\frac{18 \mathrm{~ms}}{24 \mu \mathrm{~s}}=750$. The correlator spacing of $12 \mu \mathrm{~s}$ defines the minimum sampling rate of 83.2 kHz . The correlation over the 750 lags can be performed efficiently using the Fast Fourier Transform (FFT) technique described in $[34,35]$ and implemented in a Fortran Program in $[36]$.
Presented above.


Figure 7.2-2 Probability that the Delay Estimate is Within the Indicated Bounds for Various $\mathrm{C} / \mathrm{No}^{\prime} \mathrm{s}$ and Correlator Spacing of $12 \mu \mathrm{~s}$ and $24 \mu \mathrm{~s}$.


Figure
7.2-3 Simulated and Theoretical Probability that the Delay Estimate is Within the Indicated Bounds for Various C/No's and a Correlator Spacing of $24 \mu \mathrm{~s}$.


Figure 7.2-4 Simulated and Theoretical Probability that the Delay Estimate is Within the Indicated Bounds for Various C/No's and a Correlator Spacing of $12 \mu \mathrm{~s}$.

### 8.0 OVERALL BEACON POSITIONING ACCURACY

Up to now, we discussed the various components of the subsystems which corrupt the signal and/or degrade the accuracy of the positioning technique. In this section, we derive an overall figure that combines all these effects to yield an overall positioning accuracy. There are basically two parameters that define the position accuracy which can be expected from the system : the equivalent range estimate error and the position dilution of precision (PDOP) factor. The former has been discussed in details in the previous sections for each subsystem and it can be assumed that the errors add on a root sum square(rss). Accordingly, the variance of the range error is given by :

$$
\begin{equation*}
\sigma_{\mathrm{I}}^{2}=\mathrm{c}^{2 *} \sigma_{\mathrm{t}}^{2}+\sigma_{\text {iono }}^{2}+\sigma_{\text {tropo }}^{2}+\sigma_{\text {sat }}^{2} \quad \mathrm{~m}^{2} \tag{8.0-1}
\end{equation*}
$$

Where $\sigma_{\mathrm{t}}=$ standard deviation of time delay estimate in seconds
$\sigma_{\text {iono }}=$ ionospheric delay standard deviation in meters
$\sigma_{\text {tropo }}=$ tropospheric delay standard deviation in meters
$\sigma_{\text {sat }}=$ satellite position error standard deviation in meters. It is assumed that the standard deviation of the satellite position on each axis is the same i.e. $\sigma_{\text {sat }}$.
$\mathrm{c}=$ speed of light $=2.9979 \times 10^{8} \mathrm{~m} / \mathrm{s}$
The PDOP for a single position line obtained from two satellites is derived indirectly in [38] and [39] and is given by :

$$
\begin{equation*}
\text { PDOP }=\frac{1}{4 \sin ^{2}(\theta / 2)-\left(\cos \phi_{2}-\cos \phi_{1}\right)^{2}} \tag{8.0-2}
\end{equation*}
$$

where $\theta$ and $\phi_{i}$ 's are defined as shown in Figure 8.0-1. If the coordinates of the two satellites are siven by $\left(x_{i}, y_{i}, z_{i}\right), i=1,2$ and the particular point of interest on the earth is defined by $(x, y, z)$, then we have :

$$
\begin{align*}
& \cos \theta=\frac{\left(x_{1}-x\right)\left(x_{2}-x\right)+\left(y_{1}-y\right)\left(y_{2}-y\right)+\left(z_{1}-z\right)\left(z_{2}-z\right)}{R_{1} R_{2}}  \tag{8.0-3}\\
& \cos \phi_{i}=\frac{x\left(x_{i}-x\right)+y\left(y_{i}-y\right)+z\left(z_{i}-z\right)}{R_{e} R_{i}}
\end{align*}
$$



Figure 8.0-1 PDOP Geometry
where $R_{i}$ is the range to satellite \#i and $R_{e}$ is the range from the point on the earth to its centre and they are defined as :

$$
\begin{align*}
& R_{i}=\sqrt{\left(x-x_{i}\right)^{2}+\left(y-y_{i}\right)^{2}+\left(z-z_{i}\right)^{2}}  \tag{8.0-4}\\
& R_{e}=\sqrt{x^{2}+y^{2}+z^{2}}
\end{align*}
$$

The line position accuracy is then defined as :

$$
\begin{equation*}
\sigma=(\mathrm{PDOP}) \sigma_{\mathrm{r}} \tag{8.0-5}
\end{equation*}
$$

For the time difference of arrival positioning technique, the intersection of two such positioning lines derived from three satellites defines the beacon location. It can be shown that the best accuracy is achieved when the two lines cross at square angle (see Figure 8.0-2(b)). In this Case, it is fair to assume that the position error on each line adds on a root sum square i.e. the accuracy of the position is the rss of the accuracy of the line derived from, let us say, satellites 1 and 2 and the accuracy of the line derived from satellites 1 and 3 i.e.:

$$
\begin{equation*}
\sigma=\sqrt{\left(\operatorname{PDOP}_{1,2} \sigma_{1,2}\right)^{2}+\left(\operatorname{PDOP}_{1,3} \sigma_{1,3}\right)^{2}} \quad \mathrm{~m} \tag{8.0-6}
\end{equation*}
$$

Where $\sigma_{i, j}$ is the standard deviation of the range error as defined in equation (8.0-1) for satellites \#i and Hj. Because the error made on the range estimate is independent of the satellites, it is fair to $^{\text {to }}$ assume that $\sigma_{1,2}=\sigma_{1,3}=\sigma_{2,3}=\sigma_{\mathrm{r}}$ and then we have :

$$
\begin{equation*}
\sigma=\sigma_{\mathrm{r}} \sqrt{\left(\mathrm{PDOP}_{1,2}\right)^{2}+\left(\mathrm{PDOP}_{1,3}\right)^{2}} \tag{8.0-7}
\end{equation*}
$$

or equivalently

$$
\sigma=\sigma_{\mathrm{T}} \quad \mathrm{PDOP}_{\mathrm{eq}}
$$

Where $\operatorname{PDOP}_{e q}=\sqrt{\left(\operatorname{PDOP}_{1,2}\right)^{2}+\left(\operatorname{PDOP}_{1,3}\right)^{2}}$

(a)

(b)

Figure 8.0-2 (a) Bad PDOP, (b) Good PDOP.

Figures 8.0-3 to 8.0-5 show the PDOP for various combinations of two of the three satellites located at $75^{\circ} \mathrm{W}, 105^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$ and for various latitudes. A spherical earth has been assumed to generate these figures. The figures show that the PDOP is minimum around the mid longitude between the two satellites and is symmetric around this point. They also show that the PDOP is slightly better for low latitudes than high latitudes. For high latitudes, the PDOP tends to be more constant than at low latitudes where it increases rapidly as it goes away form the mid longitude point. This is expected as the point lies in the same plane as the geostationary satellite orbit plane. Note the change of scale for Figure $8.0-5$ where it is shown that a $60^{\circ}$ satellite spacing results in a PDOP of approximately 1.

Figures 8.0-6 to 8.0-8 show the total PDOP when two position lines are combined according to equation (8.0-7). As in the other figures, a minimum is noticed and the curves are symmetric around this point. The same observations as in the previous figures can be made here. In general, it shows that the PDOP for Canada if these three satellites are used is between 1.9 and 2.8 for the best case and between 2.5 and 3.2 for the worst case.


Figure 8.0-3 PDOP for Satellites at $75^{\circ} \mathrm{W}$ and $105^{\circ} \mathrm{W}$.


Figure 8.0-4 PDOP for Satellites at $105^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$.


Figure 8.0-5 PDOP for Satellites at $75^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$.


Figure 8.0-6 Resulting Equivalent PDOP for the Combination of the Position Lines from Satellites at $75^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$ and Satellites at $75^{\circ} \mathrm{W}$ and $105^{\circ} \mathrm{W}$


Figure 8.0-7 Resulting Equivalent PDOP for the Combination of the Position Lines from Satellites at $75^{\circ} \mathrm{W}$ and $105^{\circ} \mathrm{W}$ and Satellites at $105^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$


Figure 8.0-8 Resulting Equivalent PDOP for the Combination of the Position Lines from Satellites at $75^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$ and Satellites at $105^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$

### 2.0 CIRCULAR ERROR PROBABILITY (CEP)

The accuracy of a positioning technique must always be given with its given degree of confidence. This requires in general the knowledge of the statistical distribution of the position errors which can be obtained only if the statistical distribution of each component of the total position error is known as well as their correlation factor. In practice, these distributions are not known and although one could attempt to develop such a statistical model, experimental data and straightforward approximations are usually preferred.

In [40] numerous methods to express the accuracy of positioning systems are discussed. The circular error probability (CEP) defines the radius of the circle centered on the true position which contains $50 \%$ of the points. It assumes that there are no bias errors present although in practice such bias inevitably do exist due to equipment. The CEP is a two-dimensional measure and is appropriate for the specification of the error in the system of interest.

In Section 8.0 we defined in equation (8.0-7) the equivalent position error to be given by :

$$
\begin{equation*}
\sigma=\sigma_{\mathrm{r}} \sqrt{\left(\mathrm{PDOP}_{1,2}\right)^{2}+\left(\mathrm{PDOP}_{1,3}\right)^{2}}=\sigma_{\mathrm{r}} \mathrm{PDOP}_{\mathrm{eq}} \tag{9.0-1}
\end{equation*}
$$

Which is equivalent to the distance rms (drms) error discussed in [40]. It is important to emphasize the assumptions made to obtain this equation. First, it is assumed that the error sources on each differential range are uncorrelated and have a zero mean with equal standard deviation ( $\sigma_{\mathrm{r}}$ ). Second, it is assumed that a good PDOP is obtained for the location such that the two position lines the at 90 degrees. Based on these assumptions, it is mentioned in [40] that the CEP is related to the above $\sigma$ (or drms) by :

$$
\begin{equation*}
\mathrm{CEP}=\sqrt{\ln (2)} \sigma \tag{9.0-2}
\end{equation*}
$$

given by mentioned in [40], twice the drms represents the $95 \%$ degree of confidence which is by twice equation (9.0-1) i.e. $2 \sigma$.
differen Section 8.0, the analysis did not consider the downlink (at 1.5 GHz ) effects on the signal $_{\text {at }}{ }_{\text {atial range error. In practice, the downlink effects could be estimated using a loopback }}$ $\mathrm{im}_{\text {port }}$ the master station. However, the ionospheric effects being approximately 15 times less ${ }^{\text {thu }}$ ortant at L -band than 406 MHz , it will have a small contribution to the total position error and ${ }^{\text {s can }}$ simply be taken into account or neglected.

In order to get an idea of the proposed system performance, let us look at some examples. Let us assume that the satellites are located nominally at $75^{\circ} \mathrm{W}, 105^{\circ} \mathrm{W}$ and $135^{\circ} \mathrm{W}$ and that we know their exact positions within 20 meter on each axis. From the end of Section 6.3.1, we find that the ionospheric rms error is 150 m in the uplink. Dividing this number by 15 for the downlink at 1.5 GHz , we get a rms error of 10 m on the downlink. The standard deviation of the differential range error $\left(\sigma_{\mathrm{r}}\right)$ is then given by :

$$
\begin{aligned}
\sigma_{\text {iono }} & =\sqrt{\left(\sigma_{\text {iono }}^{\text {up }}\right)^{2}+\left(\sigma_{\text {iono }}^{\text {down }}\right)^{2}} \\
& =\sqrt{150^{2}+10^{2}}=150.33 \mathrm{~m} \quad \text { (see Section 6.3.1) } \\
\sigma_{\text {sat }} & =20 \mathrm{~m} \\
\sigma_{\mathrm{t}} & =\frac{12 \mu \mathrm{~s}}{\sqrt{3}}=6.9 \mu \mathrm{~s}(\text { with } \approx 90 \% \text { confidence level at } \mathrm{C} / \mathrm{No}=30 \mathrm{~dB}-\mathrm{Hz}) \\
\sigma_{\text {tropo }} & =0 \text { (negligible relative to other sources of errors) } \\
\sigma_{\mathrm{r}}^{2} & =\mathrm{c}^{2 *} \sigma_{\mathrm{t}}^{2}+\sigma_{\text {iono }}^{2}+\sigma_{\text {tropo }}^{2}+\sigma_{\text {sat }}^{2} \quad \mathrm{~m}^{2} \\
\sigma_{\mathrm{r}} & =\sqrt{\left[\left(2.9979 \times 10^{8}\right)\left(6.9 \times 10^{-6}\right)\right]^{2}+(150.33)^{2}+(20)^{2}} \\
\sigma_{\mathrm{r}} & =2.1 \mathrm{~km}
\end{aligned}
$$

In Section 8.0 it has been seen that the (PDOP) eq was between 1.9 and 2.8 with these satellites and with a good choice of combination of satellites. The drms error and CEP are then bounded to:

$$
\sigma_{\mathrm{r}}(\mathrm{PDOP})_{\mathrm{eq}_{\min }}<\sigma<\sigma_{\mathrm{r}}(\mathrm{PDOP})_{\mathrm{eq}_{\max }}
$$

$$
4 \mathrm{~km}<\sigma<5.9 \mathrm{~km}
$$

$$
3.3 \mathrm{~km}<\text { CEP }<4.9 \mathrm{~km}
$$

provided there are no biases in the system.

It is interesting to note that the performance is dictated by the performance of the time difference of arrival estimator. In the above example, ignoring the errors introduced by the troposphere and the ionosphere would have little impact on the overall accuracy of the beacon position. This fact may be used to simplify the receiver design in an operational system.

### 10.0 RECEIVER STRUCTURE

Due to the low likelihood that three satellites will be available for use in a 406 MHz distress beacon positioning system, very little effort has been spent on the hardware design of the proposed receiver. The intent of this section is to give the reader an idea of the hardware and software required to develop the proposed distress beacon location system.

Figure 10.0-1 shows a block diagram of the receiver. The satellite receiver front ends are similar to the ones being used in the current GOES processor. It includes a low noise amplifier, a frequency down-converter to bring the signal to an intermediate frequency of 5 kHz and a satellite beacon tracker for frequency correction. Note also, it is assumed that the satellite beacon carries the satellite instantaneous position. This information is assumed to be derived by the receiver front end and it will represent a modification to the front end processor in use in the GOES project. The three satellite signals are then sampled at a rate of 83.2 kHz . Several data acquisition systems are available on the market and can be adapted to any personal computers (Macintosh and IBM compatibles). All the shaded area represents the digital processing under the control of a master controller. The detection, frequency estimation and down-conversion to baseband is already performed in the current GOES processor at a sample rate of 20.8 kHz . Accordingly, the same processing algorithms can be used here if they are preceded with a decimation by four (4). The delay estimation refers to the implementation of the correlation computation between each pair of signals. Results are sent to the master controller which interfaces to the user and monitors/controls the operation of each subsystem (e.g. if a signal is detected then freeze the memory, perform the frequency estimation and so on to come out with a position estimate). The GPS receiver is used to get the ionospheric model parameters if required.

The digital processing could be done within a personal computer (IBM 386 or compatible The Macintosh Iffx equipped with at least one array processor as well as a data acquisition system. The required number of array processors is related to the available time between events. The computer simulation which was generating the signal and noise samples in addition to computing the correlation was taking approximately 2 minutes per signal pair on a Macintosh IIcx computer equipped with a co-processor. Given that array processors provide 40-50 times more computing ${ }^{\text {Power than }}$ a general purpose computer, it is estimated that no more than $15-20$ seconds will be required to process an event i.e. a signal detection event. This assumes that data can be transferred


Figure 10.0-1 Block Diagram of Receiver and Estimator

### 11.0 CONCLUSION

This report presented a technique to locate distress beacons using geostationary satellites. In looking at the various techniques in general, it was found that a minimum of two satellites is required and that the small Doppler frequency shift of the geostationary satellites could in principle be used to resolve ambiguities. In the studied system, this information has not been used because it was believed that it would require sophisticated on-board oscillators that are unlikely going to be available on a transponder type satellite.

The time difference of arrival technique based on three satellites has been fully studied. The effect of the Earth oblateness and the ionosphere and troposphere delays have been modeled. A time difference estimator has been designed and analyzed. Theoretical performance of the estimator supported by computer simulation results demonstrated that the position error and confidence level are similar to the current system. Specifically, it can achieve a CEP between 3.3 30 and 4.9 km for beacons in Canada and a satellite spacing of $30^{\circ}$ under the worst case C/No of ${ }^{3} 0 \mathrm{~dB}-\mathrm{Hz}$ obtained when 8 distress beacons are simultaneously activated. The estimator draws a ${ }^{\text {lot }}$ from the current GOES processor in order to minimize the amount of effort required to Supplement the current distress alerting feature with distress positioning. The resources required to implement the digital processing part of the receiver are estimated to approximately 2 person-year and $\$ 75,000$ for the hardware and software.

It is obvious that the report did not address all of the techniques in great details and additional analyses should be done. In particular, a better analysis on the use of the geostationary Satellite $D_{\text {oppler }}$ frequency shift to supplement the time difference of arrival technique with only ${ }^{\text {tw }}$ o satel lites must be done. The impact of oscillator instabilities on the overall accuracy must be analyzed and traded off with other system parameters. In addition, the time difference estimator ${ }^{\text {desig }} \mathrm{ign}_{\mathrm{n}}$ should be reviewed to include the consideration of correlation in the frequency domain $\mathrm{in}_{\text {stead }}$ of the time domain. Some papers have recently pointed out that correlation in the frequency ${ }^{d} \mathrm{maxain}$ is more robust against noise than in the time domain [42].

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An extensive literature search has been conducted at the beginning of the contract. The papers are listed below under five headings :
A. GEOLOCATION TECHNIQUES AND SYSTEMS
B. COSPAS-SARSAT SYSTEM AND 406 MHZ BEACONS
C. DISTRESS SIGNAL PROCESSOR
D. PROPAGATION EFFECTS

## E. POSITION ACCURACY AND OTHER TOPICS

The papers under each heading are given in chronological order starting with the most recent.

## A. GEOLOCATION TECHNIOUES AND SYSTEMS

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## APPENDIX A

COMPUTER PROGRAM TO COMPUTE THE POSITION LINE GENERATED BY A GIVEN DIFFERENTIAL RANGE FOR TWO GIVEN SATELLITES

## LISTING OF PROGRAM 'DIFF RANGE XYZ.F'

```
ー
```

C Program to compute the line of constant range

```
C Program to compute the line of constant range
    difference given two satellites on geostationary orbits.
    difference given two satellites on geostationary orbits.
    Requires the following files :
    Requires the following files :
        RANDELT.F : to compute the ground station longitude
        RANDELT.F : to compute the ground station longitude
                                    given its latitude and its diff. range to two
                                    given its latitude and its diff. range to two
                                    given satellites
                                    given satellites
    'WGS CONVERSION.F' : to convert from WGS-84
    'WGS CONVERSION.F' : to convert from WGS-84
                                (lat., long.) coordinates to ( \(x, y, z\) )
                                (lat., long.) coordinates to ( \(x, y, z\) )
                            'Diff Range XYZ.make' : to compile and link the program
                            'Diff Range XYZ.make' : to compile and link the program
    It is strongly recommended to use the option -extended
    It is strongly recommended to use the option -extended
    during the compilation to get the maximum accuracy.
    during the compilation to get the maximum accuracy.
    NOTE : This program works fine assuming the beacon is in Canada
    NOTE : This program works fine assuming the beacon is in Canada
        and the satellites are over Canada and do not cross the
        and the satellites are over Canada and do not cross the
        180 degrees meridian
        180 degrees meridian
                    M. CARON
                    M. CARON
                    DEC. 1989
                    DEC. 1989
                            Revised Sept. 1990
                            Revised Sept. 1990
            PROGRAM DIF RANGEXYZ
            PROGRAM DIF RANGEXYZ
            IMPLICIT DOŪBLE PRECISION (A-H,L,O-Z)
            IMPLICIT DOŪBLE PRECISION (A-H,L,O-Z)
            DIMENSION SATX (2), SATY (2), SATZ (2)
            DIMENSION SATX (2), SATY (2), SATZ (2)
            CHARACTER FIIEN* 80 ,ANS*1
            CHARACTER FIIEN* 80 ,ANS*1
            C this common block passes the semi-major axis and squared eccentricity
            C this common block passes the semi-major axis and squared eccentricity
            \(C\) to the WGS CONVERSION.F routines
            \(C\) to the WGS CONVERSION.F routines
            COMMON /DISTDATA/ AE,ECC2
            COMMON /DISTDATA/ AE,ECC2
            C
            C
C Semi-major axis and squared eccentricity according to WGS-84
C Semi-major axis and squared eccentricity according to WGS-84
\(C\) npt : number of points on the line crossing the North hemisphere
\(C\) npt : number of points on the line crossing the North hemisphere
C
C
            \(A E=6378137 . D+00\)
            \(A E=6378137 . D+00\)
            \(\mathrm{ECC} 2=6.694379990 \mathrm{D}-3\)
            \(\mathrm{ECC} 2=6.694379990 \mathrm{D}-3\)
                    npt \(=100\)
                    npt \(=100\)
C
C
                    WRITE \((6,1)\)
                    WRITE \((6,1)\)
    1 FORMAT (/,/,/,T16,51('*'),/,T16,'*',49X,'*',/,T16,
    1 FORMAT (/,/,/,T16,51('*'),/,T16,'*',49X,'*',/,T16,
            \(\& \quad 1 \star\) POSITION ESTIMATION BASED ON DIFFERENTIAL RANGE *',
            \(\& \quad 1 \star\) POSITION ESTIMATION BASED ON DIFFERENTIAL RANGE *',
            \&/,T16,'*', 49X,'*',/,T16,51 ('*'),/,/,/,T5,
            \&/,T16,'*', 49X,'*',/,T16,51 ('*'),/,/,/,T5,
            \& 'NOTE : Longitudes must be within [-180,180] (+ve=East)','/
            \& 'NOTE : Longitudes must be within [-180,180] (+ve=East)','/
            \&,T5,' Latitude must be within \([-90,90]\) (+ve \(=\) North) \({ }^{\prime}\)
            \&,T5,' Latitude must be within \([-90,90]\) (+ve \(=\) North) \({ }^{\prime}\)
            \& , /, /, /)
            \& , /, /, /)
C
C
C INPUT VARIABLES
C INPUT VARIABLES
5 WRITE \((5,10)\)
5 WRITE \((5,10)\)
    10 FORMAT (/,T3,'Satellite lat., 1ong. \#1 and \#2'
```

    10 FORMAT (/,T3,'Satellite lat., 1ong. \#1 and \#2'
    ```

\section*{LISTING OF PROGRAM＇DIFF RANGE XYZ．F＇CONTINUED}
のロロのดの
```

            & ,/,T3,'Satellite #1 must be west of satellite #2')
                READ (5,*) SLAT1,SLON1, SLAT2,SLON2
                        IF ((DABS (SLON1).GT.180.D+00).OR. (DABS (SLON2).GT.180.D+00)) THEN
                    WRITE (6,*) '?? Longitudes must be within +/- 180 deg. ??'
                    GOTO }
            END IF
            IF ((DABS (SI_AT1).GT.90.D+00).OR. (DABS (SLAT2).GT.90.D+00)) THEN
                                    WRITE (6,*) ' ?? Latitudes must be within +/- 90 deg. ??'
                    GOTO 5
            END IF
    C
            20 WRITE (6,30)
            30 FORMAT (T3,'Range difference (#2 minus #1) in meters')
            READ (5,*) RDELTA
                IF(RDELTA.EQ.O.D+00) THEN
                    WRITE (6,*)'?? CANNOT BE ZERO ??'
                    GOTO 20
            END IF
    C
    40 WRITE (6,50)
    50 FORMAT (T3,'Range difference accuracy in meters')
        READ (5,*) ACC
        IF (ACC.LE.O.D+OO) THEN
                    WRITE (6,*) '?? MUST BE GREATER THAN ZERO ??'
                    GOTO 40
        END IF
    C
    51 WRITE (6,52)
    52 FORMAT (T3,'File name where to dump data')
        READ (5,*) FILEN
        OPEN (UNIT=98,FILE=FILEN,STATUS='UNKNOWN')
        REWIND (98)
        compute the satellite altitude above the earth
        approximate to geo. orbit minus earth semi-major axis
            HSAT = 42157197.D+00-AE
    C convert satellite lat, long. coordinates to (x,y,z)
CALL XYZ (SLAT1,SLON1,HSAT,SATX (1), SATY (1), SATZ (1))
CALL XYZ (SLAT2, SLON2, HSAT, SATX (2),SATY (2),SATZ (2))
WRITE (6,60)
60 FORMAT (/,/,T3,' \# ',5X,'LATITUDE',5X,'LONGITUDE', /)
C knowing that the beacon is on the north hemisphere
scan the north hemisphere in 'npt' latitude steps
flatinc = 82.d+00/float(npt-1)
WRITE (98,*) NPT
DO }100\mathrm{ I=0, npt-1

```
    \(C\)
\(C\)
\(C\)
\(C\)
C
C
            \(\mathrm{XLON}=-170 . \mathrm{D}+00\)
            XIAT \(=\) DFLOAT (I)*flatinc
C
    65 CALL RANDELT (RDELTA, 1, 2, SATX, SATY, SATZ, XIAT,
            \& \(A C C, X L O N, G L O N, I E R)\)
\(C\) if \(I E R=2\) then the subroutine could not reach the required accuracy
\(C\) with step size as small as l.d-l0 degree
\(C\) if IER=3 that means that the routine failed to find a solution
            (this should not happen unless there is erronous entries)
    if IER=0 means no error occured
\(C\) If the accuracy has not been reached, then set \(I C=1\)
C
            IF (IER.IT. 3 .AND. IER.NE. O) THEN
                    WRITE \((6, \star)\) 'IER=', IER
                    IC=1
                ELSE
                    \(I C=0\)
                END IF
                IF (IER.EQ.3) GOTO 100
C
\(C\) compute the longitude in deg. min. sec.
C
        XLON \(=\) GLON
        XDEG=DFLOAT (IDINT (DABS (XION)))
        XMIN \(=\) DFLOAT (IDINT ( (DABS (XION) \(-X D E G) * 60)\).
        \(X S E C=(D A B S(X L O N)-X D E G-X M I N / 60) * 3600.\).
        WRITE (6,70) (I+1), XLAT,XLON, XDEG,XMIN,XSEC, ic
    70 FORMAT (T3,I5,5X,F8.2,5X,F15.7,3X,F5.0,' DEG.
        \& F3.0,' min. ',F8.4,' sec.', \(2 \mathrm{x}, \mathrm{i} 2\) )
C
    75 format (f, 2x,f)
C
    100 CONTINUE
C
            WRITE \((98,62)\) SLAT1, SLON1, SLAT2, SLON2, RDELTA, ACC
    62 FORMAT (/,T3,'DIFF RANGE XYZ',/,T3,
        \& 'Sat. geocentric lat., long. \#l, \#2: ', \(2(F 7.2,1,1, F 8.2), 1\),
        \$ T3,'Range differences (meters) : ', F15.3,/,T3,
        \$ 'Accuracy Required (meters): ',F15.7.1)
C
\(C\) write an end of file to the file and close it
C
            ENDFIIE (98)
            CLOSE (98)
            WRITE \((6,200)\)
    200 FORMAT (T3,'Another run ? (Y or N)')
            READ \((5, \star)\) ANS
            IF (ANS.NE.'Y' .AND. ANS.NE.'Y') STOP 'Tourlou !'
            GOTO 5
            END

\section*{LISTING OF PROGRAM RANDELT.F}


\section*{LISTING OF PROGRAM RANDELT.F (CONTINUED)}
```

C two satellites does not cross the 180 deg. meridian line
C First convert the satellite ( $x, y, z$ ) coordinates to geodetic
C lat., long coordinates
C
CALL GEODETIC (SATX(I), SATY(I), SATZ (I), SLAT (1), SLON (1), HT)
CALL GEODETIC (SATX (J), SATY (J), SATZ (J), SLAT (2), SLON (2), HT)
c
c set IS to the west most satellite
C
$I S=I$
$I F F=J$
IF (SLON (1).GT.180.D+00) SLON(1)=SLON (1)-360.D+00
IF (SLON (2) .GT. 180.D+00) SLON (2) =SLON (2) -360.D+00
IF (SLON (1).GT.SLON (2)) THEN
IS=J
$I F F=I$
END IF
C
$C$ WRITE (6,*) 'is,iff,i,j =',IS,IFF,I,J
$C$ if satellites have been interchanged then change the sign
$C$ of the difference
IF (IS.NE.I) DELTA $=-$ DELTA
C
C set the step size to 1 degree initially
C DLON is set to 10 degrees west of the initial guess longitude
$D S T E P=1 . D+00$
DLON=XLON - 10.D+00
C
50 CONTINUE
C
GLON=DLON
C
C COMPUTE THE RANGES
$C$ First compute the ( $x, y, z$ ) coordinates of the location given
C by XLAT and DLON
C The XYZ routine works with longitudes defined between
C 0 to 360 degrees
C
TLON $=$ DLON
IF (DLON.LT. O.D+00) TLON=DLON+360.D+00
CALL XYZ (XIAT,TLON, O.D $+00, \mathrm{X}, \mathrm{Y}, \mathrm{Z}$ )
C
RI=DISTANC (SATX (IFF), SATY (IFF), SATZ (IFF), X,Y,Z)
RJ=DISTANC (SATX (IS), SATY (IS), SATZ (IS), X,Y,Z)
C
DIF2 $=\mathrm{RJ}-\mathrm{RI}$
C
C
write (6,*) 'dif2, rj, ri = ',DIF2,RJ,RI
PAUSE
C because satellite \#IFF is east of satellite \#IS and
$C$ we take $R(i s)$ - $R$ (iff) and the longitude is scanned
C from west to east, the differential range will
$C$ be negative and will gradually increase with the
$C$ increase of longitude

```

\section*{LISTING OF PROGRAM RANDELT.F (CONTINUED)}
```

        1 2 7
        118 C
        218 C
        218
        C IF (DIF2.GE.O.D+00 .AND. DELTA.GT.O.D+OO) THEN
        C If differential range greater than the required one
        then we passed the point. Come back on the previous point,
        decrease the step
        C size and increase the longitude by the new step size
        C Otherwise check if less than the required one
        C if less then increase the longitude.
        IF(DIF2.GT.(DELTA+ACC)) THEN
                        DLON=DLON-DSTEP
        DSTEP=DSTEP/2.D+00
        IF(DSTEP.LT.I.D-10) THEN
                IER=2
                RETURN
                END IF
                    DLON = DLON + DSTEP
        ELSE IF (DIF2.GE.(DELTA-ACC)) THEN
        RETURN
        ELSE
            DLON=DLON+DSTEP
        END IF
    C
    C ELSE IF (DIF2.LT.O.D+00 .AND. DELTA.GT.O.D+O0) THEN
    C if outside the allowed accuracy then increase the longitude
    C to converge toward positive DIF2
        IF (DIF2.LT.(DELTA-ACC)) THEN
        DLON = DLON + DSTEP
            ELSE
                RETURN
            END IF
                ELSE IF(DIF2.GE.O.D+00 .AND. DELTA.LT.O.D+O0) THEN
            253
            254
            155
            156
            157
            l58
            l59
            160
            l2
            162
            25
            170 C
    ```
            164
            165
            266
            167
            168
            169
            271
172
    173
274

\section*{LISTING OF PROGRAM RANDELT.F (CONTINUED)}

175
176
177
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179
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193
194
195
196
197
198
199
200

C then we passed the point. Come back on the previous point,
C decrease the step size and increase the longitude by the new
C step size
```

C

```
IF (DIF2.LT. (DELTA-ACC)) THEN
                    DLON=DLON+DSTEP
                    ELSE IF (DIF2.GT. (DELTA \(+A C C\) )) THEN
                    DLON=DLON-DSTEP
                    DSTEP=DSTEP/2.D+00
                    IF (DSTEP.LT.1.D-10) THEN
                IER=2
                                    RETURN
                    END IF
                    DLON \(=\) DLON + DSTEP
                ELSE
                    RETURN
                END IF
        END IF
C
C if longitude is greater than 180 degrees then we failed to
C find a solution
            IF (DION.LE.180.OD+OO) GOTO 50
        IER \(=3\)
        RETURN
        END

\section*{LISTING OF FILE 'WGS CONVERSION.F'}


\section*{LISTING OF FILE 'WGS CONVERSION.F' (CONTINUED)}
```

59
6 0
6 1
6 2
6 3
6 4
6 5
6 6
6 7
6 8
6 9
7 0
7 1
7 2
7 3
7 4
7 5
7 6
7 7
7 8
7 9
80
81
82
8
84
8
8
8
8
89
90
91
92
93
94
95
96
97
a2=e2/2.D0

```
```

        a3=a2/2.D0
        a4=al**2.DO/2.DO
        a5=al/2.DO
        a6=1.D0-e2
        xln=datan2 (y,x)
        w2=x**2.DO+y**2.DO
        w=dsqrt (w2)
        r2=w2+z**2.D0
        r=dsqrt(r2)
        s22=z**2.DO/r2
        d1=al/r
        d2=d1+1.D0
        c=a5*s22* (d2+s22* (a3-d1))
        ht=r-ae+c
        s2=z/r
        c2=w/r
        sl=d1*s2*c2*(d2-s22* (2.D0*d1-a.a))
        cl=1.DO-(a.4/r2)*s22*(1.DO-s22)
        s=s1*c2+cl*s2
        ylt=dasin(s)
        ss=s**2.DO
    rr=1.D0-e2*ss
    re=ae/dsqrt (rr)
    rf=a.6*re
    c=dsqrt (1.D0-ss)
    wl= (re+ht)*c
    zl=(rf+ht)*s
    dw=w-wl
    dz=z-zl
    dl=-s*dw+C*dz
    dh=c*dw+s*dz
    ylt=ylt+dl/(rf/rr+ht)
    ht=ht+dh
        dtor = datan(1.0d+0)/45.d+00
        ylt = ylt / dtor
        xln = xln / dtor
    return
    end
    ```

\section*{LISTING OF SCRIPT FILE TO COMPILE AND LINK}
\# File: 'Diff Range XYZ.make'
\# Target: 'Diff Range XYZ'
\# Sources: 'DIFF RANGE XYZ.F' randelt.f 'WGS84 CONVERSION.F'
\# Created: Friday, February 2, 1990 10:27:15 AM
'DIFF RANGE XYZ.F.o' \(f\) 'Diff Range XYZ.make' 'DIFF RANGE XYZ.F' FORTRAN - 3 -mc 68881 -opt \(=3\)-extended 'DIFF RANGE XYZ.F'
randelt.f.o \(f\) 'Diff Range XYZ.make' randelt.f
FORTRAN -3 -mc68881 -opt=3 -extended randelt.f
'WGS84 CONVERSION.F.o' \(f\) 'Diff Range XYZ.make' 'WGS84 CONVERSION.F' FORTRAN -3 -mc68881 -opt=3 -extended 'WGS84 CONVERSION.F'

SOURCES = 'DIFF RANGE XYZ.F' randelt.f 'WGS84 CONVERSION.F'
OBJECTS = 'DIFF RANGE XYZ.F.o' randelt.f.o 'WGS84 CONVERSION.F.o'
'Diff Range XYZ' ff 'Diff Range XYZ.make' \{OBJECTS \}
Link -ad 4 -w -t APPL -c '????' \(\partial\)
\{OBJECTS \} \(\partial\)
"\{Libraries\}"Runtime.o \(\partial\)
"\{Libraries \}"Interface.o д
" (FLibraries \}"FORTRANlib.o \(\partial\)
"\{FLibraries\}"IntrinsicLib.o д
-o 'Diff Range XYZ'

\section*{RUN EXAMPLE}
```

********************************************************

*     * 
* POSITION ESTIMATION BASED ON DIFFERENTIAL RANGE *
* 

********************************************************

```

NOTE : Longitudes must be within [-180,180] (+ve=East)
Latitude must be within \([-90,90]\) (+ve \(=\) North \()\)

Satellite lat., long. \#1 and \#2
Satellite \#1 must be west of satellite \#2
0,-135,0,-75
Range difference (\#2 minus \#1) in meters
215000
Range difference accuracy in meters 1
File name where to dump data demo.dat
\# LATITUDE LONGITUDE
\begin{tabular}{|c|c|c|c|c|}
\hline 1 & 0.00 & -103.3149719 & 103. DEG. & n. 53.8989 sec .0 \\
\hline 2 & 0.83 & -103.3147736 & 103. DEG. & 18. min. 53.1848 sec .0 \\
\hline 3 & 1.66 & -103.3141479 & 103. DEG. & 18. min. 50.9326 sec .0 \\
\hline 4 & 2.48 & -103.3131256 & 103. DEG. & 18. min. 47.2522 sec .0 \\
\hline 5 & 3.31 & -103.3116913 & 103. DEG. & 18. min. 42.0886 sec .0 \\
\hline 6 & 4.14 & -103.3098297 & 103. DEG. & 18. min. 35.3870 sec .0 \\
\hline 7 & 4.97 & -103.3075714 & 103. DEG. & 18. min. 27.2571 sec .0 \\
\hline 8 & 5.80 & -103.3048859 & 103. DEG. & 18. min. 17.5891 sec .0 \\
\hline 9 & 6.63 & -103.3017731 & 103. DEG. & 18. min. 6.3831 sec 0 \\
\hline 10 & 7.45 & -103.2982483 & 103. DEG. & 17. min. 53.6938 sec . 0 \\
\hline 11 & 8.28 & -103.2942810 & 103. DEG. & 17. min. 39.4116 sec . \\
\hline 12 & 9.11 & -103.2899017 & 103. DEG. & 17. min. 23.6462 sec . \\
\hline 13 & 9.94 & -103.2850800 & 103. DEG. & 17. min. 6.2878 sec 0 \\
\hline 14 & 10.77 & -103.2798004 & 103. DEG. & 16. min. 47.2815 sec . \\
\hline 15 & 11.60 & -103.2740936 & 103. DEG. & 16. min. 26.7371 sec . \\
\hline 16 & 12.42 & -103.2679291 & 103. DEG. & 16. min. 4.5447 sec . \\
\hline 17 & 13.25 & -103.2612915 & 103. DEG. & 15. min. 40.6494 sec . \\
\hline 18 & 14.08 & -103.2541962 & 103. DEG. & 15. min. 15.1062 sec . \\
\hline 19 & 14.91 & -103.2466278 & 103. DEG. & 14. min. 47.8601 sec . \\
\hline 20 & 15.74 & -103.2385712 & 103. DEG. & 14. min. 18.8562 sec . \\
\hline 21 & 16.57 & -103.2300262 & 103. DEG. & 13. min. 48.0945 sec . \\
\hline 22 & 17.39 & -103.2209778 & 103. DEG. & 13. min. 15.5200 sec . \\
\hline 23 & 18.22 & -103.2114105 & 103. DEG. & 12. min. 41.0779 sec . \\
\hline 24 & 19.05 & -103.2013245 & 103. DEG. & 12. min. 4.7681 sec . \\
\hline 25 & 19.88 & -103.1907043 & 103. DEG. & 11. min. 26.5356 sec . \\
\hline
\end{tabular}

\section*{RUN EXAMPLE (CONTINUED)}


\section*{RUN EXAMPLE (CONTINUED)}


STOP Tourlou !

\section*{RUN EXAMPLE (CONTINUED)}

100
0.000000000000000000
0.828282828282828283
1.656565656565656566
2.484848484848484848
3.313131313131313131
4.141414141414141414
4.969696969696969696
5.797979797979797980
6.626262626262626262
7.454545454545454546
8.282828282828282828
9.111111111111111112
9.939393939393939393
10.767676767676767680
11.595959595959595960
12.424242424242424240
13.252525252525252530
14.080808080808080810
14.909090909090909090
15.737373737373737370
16.565656565656565660
17.393939393939393940
18.22222222222222220
19.050505050505050500
19.878787878787878790
20.707070707070707070
21.535353535353535350
22.363636363636363640
23.191919191919191920
24.020202020202020200
24.848484848484848480
25.676767676767676770
26.505050505050505050
27.333333333333333330
28.161616161616161620
28.989898989898989900
29.818181818181818180
30.646464646464646460
31.474747474747474750
32.303030303030303030
33.131313131313131310
33.959595959595959600
34.787878787878787880
35.616161616161616160
36.444444444444444440
37.272727272727272730
38.101010101010101010
38.929292929292929290
39.757575757575757580
40.585858585858585860
41.414141414141414140
42.242424242424242420
43.070707070707070710
-103.314971923828125000
\(-103.314773559570312500\)
-103.314147949218750000
\(-103.313125610351562500\)
-103.311691284179687500
-103.309829711914062500
-103.307571411132812500
-103.304885864257812500
\(-103.301773071289062500\)
\(-103.298248291015625000\)
-103.2942810058559375000
-103.289901733398437500
-103.285079956054687500
-103.279800415039062500
-103.274093627929687500
-103.267929077148437500
-103.261291503906250000
-103.254196166992187500
-103.246627807617187500
-103.238571166992187500
-103.230026245117187500
-103.220977783203125000
-103.211410522460937500
-103.201324462890625000
-103.190704345703125000
-103.179519653320312500
-103.167785644531250000
-103.155471801757812500
-103.142578125000000000
-103.129058837890625000
-103.114929199218750000
-103.100143432617187500
\(-103.084686279296875000\)
-103.068572998046875000
-103.051727294921875000
-103.034164428710937500
-103.015853881835937500
-102.996749877929687500
-102.976852416992187500
-102.956115722656250000
-102.934494018554687500
-102.911987304687500000
-102.888519287109375000
-102.864089965820312500
-102.838653564453125000
\(-102.812133789062500000\)
-102.784515380859375000
-102.755752563476562500
-102.725769042968750000
\(-102.694519042968750000\)
\(-102.661956787109375000\)
\(-102.627990722656250000\)
-102.592559814453125000

RUN EXAMPLE (CONTINUED)
\begin{tabular}{ll}
43.898989898989898990 & -102.555603027343750000 \\
44.727272727272727270 & -102.517028808593750000 \\
45.555555555555555560 & -102.476745605468750000 \\
46.383838383838383840 & -102.434692382812500000 \\
47.212121212121212120 & -102.390716552734375000 \\
48.040404040404040400 & -102.344757080078125000 \\
48.868686868686868680 & -102.296661376953125000 \\
49.696969696969696960 & -102.246337890625000000 \\
50.525252525252525260 & -102.193603515625000000 \\
51.353535353535353540 & -102.138366699218750000 \\
52.181818181818181820 & -102.080413818359375000 \\
53.010101010101010100 & -102.019561767578125000 \\
53.838383838383838380 & -101.955657958984375000 \\
54.666666666666666660 & -101.888488769531250000 \\
55.494949494949494940 & -101.817779541015625000 \\
56.323232323232323240 & -101.743316650390625000 \\
57.151515151515151520 & -101.664794921875000000 \\
57.979797979797979800 & -101.581909179687500000 \\
58.808080808080808080 & -101.494293212890625000 \\
59.636363636363636360 & -101.401580810546875000 \\
60.464646464646464640 & -101.303344726562500000 \\
61.292929292929292920 & -101.199127197265625000 \\
62.121212121212121220 & -101.088378906250000000 \\
62.949494949494949500 & -100.970489501953125000 \\
63.777777777777777780 & -100.844787597656250000 \\
64.606060606060606060 & -100.710540771484375000 \\
65.434343434343434340 & -100.566833496093750000 \\
66.262626262626262620 & -100.412719726562500000 \\
67.090909090909090900 & -100.247070312500000000 \\
67.919191919191919200 & -100.068542480468750000 \\
68.747474747474747480 & -99.875640869140625000 \\
69.575757575757575760 & -99.666625976562500000 \\
70.404040404040404040 & -99.439422607421875000 \\
71.232323232323232320 & -99.191650390625000000 \\
72.060606060606060600 & -98.920349121093750000 \\
72.888888888888888880 & -98.622192382812500000 \\
73.717171717171717160 & -98.292968750000000000 \\
74.545454545454545460 & -97.927734375000000000 \\
75.373737373737373740 & -97.520141601562500000 \\
76.202020202020202020 & -97.062622070312500000 \\
77.030303030303030300 & -96.545410156250000000 \\
77.858585858585858580 & -95.956115722656250000 \\
78.686868686868686860 & -95.278625488281250000 \\
79.515151515151515160 & -94.491577148437500000 \\
80.343434343434343440 & -93.566040039062500000 \\
81.171717171777171720 & -92.461791992187500000 \\
82.000000000000000000 & -91.121582031250000000 \\
\hline
\end{tabular}

\section*{DIFF RANGE XYZ}

Sat. geocentric lat., long. \#1, \#2 : \(0.00,-135.00 \quad 0.00,-75.00\) Range differences (meters) : \(\quad 215000.000\) Accuracy Required (meters) : 1.0000000

\section*{APPENDIX B}

COMPUTER PROGRAM TO COMPUTE THE SHORTEST DISTANCE BETWEEN TWO POINTS ON THE EARTH

\section*{LISTING OF PROGRAM DISVINCENTY.F}
```

program disvincenty

```
```

c This program calculates geodetic distance using Vincenty's Inverse

```
c This program calculates geodetic distance using Vincenty's Inverse
c algorithm. Based on paper presented at 39th annual meeting of ION
c algorithm. Based on paper presented at 39th annual meeting of ION
c by L. Pfeifer pp515-524.
c by L. Pfeifer pp515-524.
c
c
c written by D. Hindson
c written by D. Hindson
c May 11990
c May 11990
implicit double precision (a-z)
implicit double precision (a-z)
integer ans,fdeg,fmin,bdeg,bmin
integer ans,fdeg,fmin,bdeg,bmin
character*1 msym,ssym
character*1 msym,ssym
msym=char(39)
msym=char(39)
    ssym=char(34)
    ssym=char(34)
    pi=4.d0*datan(1.d0)
    pi=4.d0*datan(1.d0)
    dtor=pi/180.d0
    dtor=pi/180.d0
    write(6,*)'Do you want to use WGS84 reference ellipsoid (y=1)'
    write(6,*)'Do you want to use WGS84 reference ellipsoid (y=1)'
    read(5,*) ans
    read(5,*) ans
    if (ans.eq.1) then
    if (ans.eq.1) then
c
c
c WGS84 parameters
c WGS84 parameters
c
c
    a=6378137.d0
    a=6378137.d0
    finv=298.257223563d0
    finv=298.257223563d0
    else
    else
            write(6,*)'Enter ellipsoid semi-major axis (m)'
            write(6,*)'Enter ellipsoid semi-major axis (m)'
            read(5,*) a
            read(5,*) a
            write(6,*)'Enter 1/flattening of ellipsoid'
            write(6,*)'Enter 1/flattening of ellipsoid'
            read(5,*) finv
            read(5,*) finv
    endif
    endif
c
c
c Input section
c Input section
c
c
300 write(6,*)'Do you want to enter lat. longs. in degrees (1) or d m s (2)'
300 write(6,*)'Do you want to enter lat. longs. in degrees (1) or d m s (2)'
        read(5,*) ans
        read(5,*) ans
        if (ans.eq.1) then
        if (ans.eq.1) then
            write(6,*)'Input start latitude (deg. N +ve)'
            write(6,*)'Input start latitude (deg. N +ve)'
            read(5,*) lats
            read(5,*) lats
            write(6,*)'Input start longitude (deg. E +ve)'
            write(6,*)'Input start longitude (deg. E +ve)'
            read(5,*) longs
            read(5,*) longs
            write(6,*)'Input finish latitude (deg. N +ve)'
            write(6,*)'Input finish latitude (deg. N +ve)'
            read(5,*) latf
            read(5,*) latf
                write(6,*)'Input finish longitude (deg. E +ve)'
                write(6,*)'Input finish longitude (deg. E +ve)'
                read(5,*) longf
                read(5,*) longf
            else
            else
                write(6,*)'Input start latitude (deg. min. sec. N +ve)'
                write(6,*)'Input start latitude (deg. min. sec. N +ve)'
                read(5,*) lats,min,sec
                read(5,*) lats,min,sec
                if (lats.lt.0.d0) then
                if (lats.lt.0.d0) then
                lats=lats-min/60.d0-sec/3600.d0
                lats=lats-min/60.d0-sec/3600.d0
            else
            else
                lats=lats+min/60.d0+sec/3600.d0
                lats=lats+min/60.d0+sec/3600.d0
                endif
```

                endif
    ```

\section*{LISTING OF PROGRAM DISVINCENTY.F (CONTINUED)}
            write \((6, *)^{\prime}\) 'Input start longitude (deg. min. sec. E +ve)'
            read (5,*) longs,min,sec
            if (longs.lt.0.d0) then
                longs \(=\) longs \(-\mathrm{min} / 60 . \mathrm{d} 0-\mathrm{sec} / 3600 . \mathrm{d} 0\)
            else
                longs \(=\) longs \(+\min / 60 . d 0+\mathrm{sec} / 3600 . \mathrm{d} 0\)
            endif
            write( \(6, *\) )'Input finish latitude (deg. min. sec. \(\mathrm{N}+\mathrm{ve}\) )'
            read(5,*) latf,min,sec
            if (latf.It.0.d0) then
                latf \(=\) latf-min \(/ 60 . \mathrm{d} 0-\mathrm{sec} / 3600 \mathrm{~d} 0\)
            else
                latf \(=\mathrm{Jatf}+\mathrm{min} / 60 . \mathrm{d} 0+\mathrm{sec} / 3600 . \mathrm{d} 0\)
            endif
            write( \(6,{ }^{*}\) )'Input finish longitude (deg. min. sec. \(\mathrm{E}+\mathrm{ve}\) )'
            \(\operatorname{read}(5, *)\) longf,min,sec
            if (longf.lt.0.d0) then
                longf=longf-min/60.d0-sec/3600.d0
            else
                longf \(=\) longf \(+\mathrm{min} / 60 . \mathrm{d} 0+\mathrm{sec} / 3600 . \mathrm{d} 0\)
            endif
        endif
            c
            c convert lat long to radians
            c note: algorithm uses \(\mathrm{W}+\mathrm{ve}\) thud -ve sign on long
            c
                lats=lats*dtor
            longs \(=-\) longs*dtor
            latf=latf*dtor
            longf \(=-\) longf*dtor
            c
                    c call subroutine dis to calc distance and azimuths
                    call dis(a,finv,lats,longs,latf,longf,faz,baz,dist)
            call d_dms(faz,fdeg,fmin,fsec)
            call d_dms(baz,bdeg,bmin,bsec)
            write (6,*)' '
            write \((6, *)^{\prime}\) Distance between points \(=\) ',dist,' \(\mathrm{m}^{\prime}\)
            write (6,*)' '
            write(6,*)' Forward azimuth = ',faz,' degrees'
            write(6,110),fdeg,fmin,msym,fsec,ssym
            write(6,*)'
            write(6,*)' Back azimuth = ',baz,' degrees'
            write (6,110), bdeg,bmin,msym,bsec,ssym
            110 format \((26 x,=', 4 x, \mathrm{i} 3,1 \mathrm{x}, \mathrm{i} 2, \mathrm{al}, \mathrm{f} 7.3, \mathrm{a} 1)\)
            write \((6, *)^{\prime}\)
            write( \(6, *)^{\prime}\) '
            write \((6, *)^{\prime}\) Do you want to run again? \((y=1)^{\prime}\)
            \(\operatorname{read}(5, *)\) ans
            if (ans.eq.1) goto 300
            stop
            end
subroutine dis(a,finv,glat1,glon1,glat2,glon2,faz,baz,s)
Implicit double precision (a-z)
integer i,ans
tol \(=0.3 \mathrm{~d}-11\)
twopi \(=6.283185307179586 \mathrm{~d} 0\)
\(\mathrm{r}=-1 . \mathrm{d} 0 / \mathrm{finv}+1 . \mathrm{d} 0\)
tul \(=\) dtan (glat 1 )* \({ }^{*}\)
tu2 \(=\) dtan (glat2)* \({ }^{\text {r }}\)
\(\mathrm{cu}=1 . \mathrm{d} 0 / \mathrm{dsqrt}\left(\mathrm{tu} 1^{*} \mathrm{tu} 1+1 . \mathrm{d} 0\right)\)
sul \(=c u 1^{*}\) tul
cu \(2=1 . \mathrm{d} 0 / \mathrm{dsqrt}(\mathrm{tu} 2 *\) tu \(2+1 . \mathrm{d} 0\) )
\(\mathrm{s}=\mathrm{cu} \mathbf{2}^{*} \mathrm{cu} 1\)
baz=s*tu2
faz=baz*tul
\(\mathrm{x}=\) glon2-glon 1
\(i=0\)
\(100 \quad \mathrm{i}=\mathrm{i}+1\)
if(i.gt.1000) then
                                    write \((6, *)^{\prime} 1000\) itterations performed continue? \((y=1)^{\prime}\)
                                    \(\operatorname{read}\left(5,{ }^{*}\right)\) ans
            if (ans.ne.1) return
endif
\(\mathrm{sx}=\mathrm{d} \sin (\mathrm{x})\)
\(\mathrm{cx}=\mathrm{d} \cos (\mathrm{x})\)
tu1 \(=\mathrm{cu} 2^{*} \mathrm{sx}\)
tu2 \(=\)-su1* cu \(2 * \mathrm{cx}+\mathrm{baz}\)
    sy=dsqrt(tu1*tu1+tu2*tu2)
    cy=s*cx+faz
    \(\mathrm{y}=\) datan2(sy,cy)
    \(\mathrm{sa}=\mathrm{s}^{*} \mathrm{~s} \mathrm{x} / \mathrm{sy}\)
    c2a \(=-s a^{*} \mathrm{sa}+1 . \mathrm{d} 0\)
    cz=faz+faz
    if( c2a.gt.0.d0) cz=-cz/c2a+cy
    \(\mathrm{e}=c z^{*} c z^{*} 2 . \mathrm{d} 0-1 . \mathrm{d} 0\)
    \(c=((-3 . d 0 * c 2 a+4 . d 0) / f i n v+4 . d 0) * c 2 a / f i n v / 16 . d 0\)
    \(\mathrm{d}=\mathrm{x}\)
    \(\mathrm{x}=\left(\left(\mathrm{e}^{*} \mathrm{cy*} \mathrm{c}+\mathrm{cz}\right)^{*} \mathrm{sy}^{*} \mathrm{c}+\mathrm{y}\right) *\) sa
    \(x=(1 . d 0-c)^{*} x /\) finv + glon2-glon 1
    if(dabs(d-x).gt.tol*y) goto 100
    faz=datan2(-tu1,tu2)
    if(faz.lt.0.d0) faz=faz+twopi
    baz=datan2(cu1*sx,su1*cu2-baz*cx)
    if(baz.lt.0.d0) baz=baz+twopi
    \(\mathrm{x}=\mathrm{dsqrt}((1 . \mathrm{d} 0 / \mathrm{r} / \mathrm{r}-1 . \mathrm{d} 0) * \mathrm{c} 2 \mathrm{a}+1 . \mathrm{d} 0)+1 . \mathrm{d} 0\)
    \(\mathrm{x}=(\mathrm{x}-2 . \mathrm{d} 0) / \mathrm{x}\)
    \(\mathrm{c}=1 . \mathrm{d} 0-\mathrm{x}\)
    \(c=(x * x / 4 . d 0+1 . d 0) / c\)
    \(\mathrm{d}=\left(0.375 \mathrm{~d} 0 *{ }^{*} * \mathrm{x}-1 . \mathrm{d} 0\right) * \mathrm{x}\)
    \(\mathrm{x}=\mathrm{e}^{*} \mathrm{cy}\)
    \(s=1\).d0-e-e
    \(s=\left(\left(\left(\left(s y^{*} s y * 4 . d 0-3 . d 0\right) * s * c z * d / 6 . d 0-x\right) * d / 4 . d 0+c z\right) * s y * d+y\right) * c^{*} a^{*} r\)
    return
    end

\section*{LISTING OF PROGRAM DISVINCENTY.F (CONTINUED)}

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\section*{subroutine d_dms(angle,deg,min,sec)}
c
c c routine convert angle in radians to decimal degrees and degrees min sec implicit double precision (a-z) integer deg,min pi=4.d0*datan(1.d0) angle=angle*180.d0/pi \(\mathrm{deg}=\) int (angle) min=int((angle-dfloat(deg))*60.d0) \(\mathrm{sec}=(\) angle-dfloat(deg)-dfloat(min)\(/ 60 . \mathrm{d} 0) * 3600 . \mathrm{d} 0\) return
end
```


# File: Disvincenty.make

# Target: Disvincenty

# Sources: disvincenty.f

# Created: Tuesday, October 2, 1990 4:15:27 PM

OBJECTS = disvincenty.f.o
Disvincenty ff Disvincenty.make {OBJECTS}
Link -f -srt -ad 4 -ss 1000000 -w -t APPL -c '????' \partial
{OBJECTS} д
"{Libraries}"Runtime.o\partial
"{Libraries}"Interface.oд
"{FLibraries}"FORTRANlib.o \partial
"{FLibraries}"IntrinsicLibFPU.o \partial
"{FLibraries}"FSANELibFPU.oд
-o Disvincenty
disvincenty.f.o f Disvincenty.make disvincenty.f
FORTRAN -mc68020 -mc68881 -opt=3 -extended disvincenty.f

```

\section*{RUN EXAMPLE}

1 Do you want to use WGS84 reference ellipsoid ( \(\mathrm{y}=1\) )
2 1
3 Do you want to enter lat. longs. in degrees (1) or d m s (2)
41
5 Input start latitude (deg. \(\mathrm{N}+\mathrm{ve}\) )
646
7 Input start longitude (deg. \(\mathrm{E}+\mathrm{ve}\) )
8 -66
9 Input finish latitude (deg. \(\mathrm{N}+\mathrm{ve}\) )
1045
11 Input finish longitude (deg. \(\mathrm{E}+\mathrm{ve}\) )
12 -75
13
14 Distance between points \(=711748.631389547865 \mathrm{~m}\)
\[
\text { Forward azimuth }=264.252737733617996 \text { degrees }
\]
\[
=26415^{\prime} 9.856^{\prime \prime}
\]

Back azimuth \(=77.8267293949662975\) degrees \(=7749^{\prime} 36.226^{\prime \prime}\)

Do you want to run again? \((\mathrm{y}=1)\)
N

STOP
.

\section*{APPENDIX C}

PROGRAM TO COMPUTE THE ACCURACY OF THE DELAY ESTIMATOR
```

C
C PROGRAM TO COMPUTE THE PROBABLITTY THAT A GIVEN ACCURACY
C IS REACHED IN THE DELAY ESTIMATOR PROCESSOR
C
M. Caron
AUGUST 1990
Note : It is suggested to use the "-extended" option when
compiling to obtain better accuracy on the calculation
RunMacII Accuracy4.f -extended -opt=3
Only the first 512 correlation lags are assumed to be
of interest.
C Load trap code for Macintosh system routines
!!m inlines.f
PROGRAM ACCURACY4
real*8 CORREL(0:512)
DOUBLE PRECISION SUM,PROD,FCN,U,sigma,AA,Q
EXTERNAL FCN,Q
CHARACTER*80 FILEN,out**80
C
do i=0,512
correl(I)=0.
end do
c
WRITE (6,10)
10 FORMAT (/, /,,T10,'DELAY ESTIMATION ACCURACY',/T10,
\& '=======================',N/()
C
50 WRITE (6,60)
60 FORMAT (T3,'FILE NAME WHERE TO GET THE',
\& 'AUTOCORRELATION LEVELS')
OPEN (UNIT=10,FILE=*,STATUS='OLD',readonly)
inquire(unit=10,name=filen)
write(6,*) filen
c
call f_drawoutpwindow
C
C read autocorrelation levels over the first NCOR lags
C
READ (10,20) NCOR
20 FORMAT(i)
IF(NCOR.GT.513) THEN
WRITE(6,25) NCOR
25 FORMAT(T3,'?? NCOR GREATER THAN THE MAXIMUM LIMIT.',
\&
'NCOR=',16,'??')
STOP
END IF
READ (10,28)(IA,CORREL(I),I=0,NCOR-1)
28 FORMAT (i,f)

```

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56
c

C

\section*{CLOSE(10)}

\section*{C}
```

80 WRITE $(6,85)$
85 FORMAT (T3,'C/NO IN DB-HZ')
READ (5,*) CNO
C
120 WRITE $(6,130)$
130 FORMAT(T3,'Enter the estimate error in number of correlators') READ (5,*) MDEL IF(MDEL.LT.0 .OR. MDEL.GT.NCOR) GOTO 120
c
182 CONTINUE
C
$S U M=0.0 \mathrm{D}+00$
$\lim =\operatorname{MIN}($ ncor,512)/10
c
c compute variance
SIGMASQ $=10 . d+00 * *(-\mathrm{CNO} / 10 . d+00) * 2 . d+00 / 0.44 d+00$
write $(6, *)$ 'sigmasq $=$ ', sigmasq
c
c assume that the response is symmerric around the real delay
c so do from 1 to mdel and multiply the prob. by 2 for the
c two side effect
c
DO $300 \mathrm{KMI}=1, \mathrm{MIN}$ (MDEL,NCOR-1)
C
PROD $=1.0 \mathrm{D}+00$
DO $200 \mathrm{KMJ}=0, \mathrm{MIN}(\mathrm{NCOR}-1,511)$
IF(KMJ.EQ.KMI) GOTO 200
SIGMA $=$ DSQRT(SIGMASQ*(1.D+00-CORREL(IABS(KMI-KMJ)))) $\mathrm{U}=-(\mathrm{CORREL}(\mathrm{KMI})-\mathrm{CORREL}(\mathrm{KMJ})) /$ SIGMA
$\mathrm{AA}=\mathrm{Q}$ (U)
PROD=PROD*AA
$\mathrm{X} \quad$ WRITE $(6,185) \mathrm{U}, \mathrm{Q}(\mathrm{U}), \mathrm{PROD}$
X185
C
200
CONTINUE
write $(6,205)$ KMI,PROD*100.
205 FORMAT (T3,'Prob. that lag \#',I2,' is the highest is ',f10.4,' \%')
SUM=SUM+PROD WRITE $(6,210) 200 . \mathrm{D}+00^{*}$ SUM
210 FORMAT(T3,'Cumulative Probability $=$ =',f10.4,' \%',/)
c if probability that the output KMI samples away from the true output
c is less than $10^{* *}(-5)$, then the prob that the output KMI +n samples
c away from the true output is negligible ( $\mathrm{n}>0$ )
c
IF(PROD.LT.1.D-5) LEAVE
300 CONTINUE
X PAUSE
$\mathrm{KMI}=0$

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$$
\text { PROD }=1.0 \mathrm{D}+00
$$

```
```

    DO 400 KMJ=1,MIN(NCOR-1,512)
    ```
    DO 400 KMJ=1,MIN(NCOR-1,512)
        SIGMA = DSQRT(SIGMASQ*(1.D+00-CORREL(LABS(KMI-KMJ))))
        SIGMA = DSQRT(SIGMASQ*(1.D+00-CORREL(LABS(KMI-KMJ))))
        U=-(CORREL(KMI)-CORREL(KMJ))/SIGMA
        U=-(CORREL(KMI)-CORREL(KMJ))/SIGMA
        AA=Q(U)
        AA=Q(U)
        PROD=PROD*AA
        PROD=PROD*AA
X WRITE (6,185) U,Q(U),PROD
X WRITE (6,185) U,Q(U),PROD
400 CONTINUE
400 CONTINUE
            write (6,205) KMI,PROD*100.
            write (6,205) KMI,PROD*100.
        SUM=2.D+00*SUM+PROD
        SUM=2.D+00*SUM+PROD
        WRITE (6,210) SUM*100.d+00
        WRITE (6,210) SUM*100.d+00
C
            WRITE (6,*) '1 FOR MORE '
            WRITE (6,*) '1 FOR MORE '
            READ(5,*) I
            READ(5,*) I
            IF(I.EQ.1) GOTO }8
            IF(I.EQ.1) GOTO }8
    STOP
    STOP
    END
    END
C
c function to compute the Q-function
c function to compute the Q-function
c
c program derived from the TI-58 pocket calculator
c program derived from the TI-58 pocket calculator
c internal program
c internal program
c
c
c M. Caron June }199
c M. Caron June }199
c
c
c
c
                double precision function Q(x)
                double precision function Q(x)
                implicit double precision (a-h,o-z)
                implicit double precision (a-h,o-z)
c
c
                    data al,a2,a3,a4,a5,a6 /0.2316419,1.330274429,1.821255978,
                    data al,a2,a3,a4,a5,a6 /0.2316419,1.330274429,1.821255978,
    & 1.781477937,0.356563782,0.31938153/
    & 1.781477937,0.356563782,0.31938153/
            DATA PI/3.14159265358979324/
            DATA PI/3.14159265358979324/
c
c
            y=x
            y=x
            minus = 0
            minus = 0
            if(y.le.0.d+00) then
            if(y.le.0.d+00) then
            minus =1
            minus =1
            y=dabs(y)
            y=dabs(y)
            end if
            end if
            c
            c
            deno = al*y + 1
            deno = al*y + 1
            b = a2/deno**4 - a3/deno**3 + a4/deno**2 - a5/deno +a6
            b = a2/deno**4 - a3/deno**3 + a4/deno**2 - a5/deno +a6
            b=b/deno/dsqrt(2.*pi)*DSQRT(exp(-(y**2)))
            b=b/deno/dsqrt(2.*pi)*DSQRT(exp(-(y**2)))
            if(minus.eq.0) then
            if(minus.eq.0) then
            Q=b
            Q=b
            else
            else
            Q=1.d+00-b
            Q=1.d+00-b
            end if
            end if
            return
            return
            end
```

            end
    ```

1

\section*{DELAY ESTIMATION ACCURACY}

FILE NAME WHERE TO GET THE AUTOCORRELATION LEVELS correl200.dat

\section*{C/NO IN DB-HZ}

30
Enter the estimate error in number of correlators
3
sigmasq \(=4.545454545454545454 \mathrm{E}-003\)
Prob. that lag \# 1 is the highest is 11.3654 \%
Cumulative Probability \(=22.7307 \%\)
Prob. that lag \# 2 is the highest is 1.0957 \%
Cumulative Probability \(=24.9222 \%\)
Prob. that lag \# 3 is the highest is \(0.0603 \%\)
Cumulative Probability \(=25.0429 \%\)
Prob. that lag \# 0 is the highest is \(64.8936 \%\)
Cumulative Probability \(=89.9365 \%\)
\({ }_{1}\) FOR MORE
2
STOP

\section*{APPENDIX D}

COMPUTER PROGRAM TO SIMULATE

\section*{THE DELAY ESTIMATION PROCESSOR}

\author{
Listing of Program 'Delay Estimator.f'
}


\section*{Listing of Program 'Delay estimator.f' (Continued)}

C
REAL*4 XMEAN,SIGMA1,SIGMA2,R1,R2,Q1,Q2
DATA XMEAN /0.0/
C
SEED1 and SEED2 are the two seeds for the random number generator
SEED1 \(=4\)
SEED2 \(=3141592\)
C
C validate some of the parameter values that couldn't be inter-related
C in the above PARAMETER statements
C compute the maximum number of lags of interest (LAGSHOW)
\({ }_{C}^{C}\) Compute the number of lags to be computed (power of two)
A = DELMX*FSMUL
LAGSHOW = IFIX(A)
IF(FLOAT(LAGSHOW).NE.A) LAGSHOW = LAGSHOW +1
I=ALOG(A)/ALOG(2.)
IF(FLOAT(2**I).NE.A) I =I + 1
LAG \(=2 * * I+1\)
IF(LAG.GT.LAGMX) STOP 'DELMX IS TOO LARGE FOR LAGMX. ABORT'
C
C compute the size of the FFT to be used
C compute the number of noise samples required
C compute the total number of samples required to represent the signal
C
C NSIZE \(=2 * *(I+1)\)
NOISE \(=\operatorname{IFIX}(\mathrm{A})\)
IF(FLOAT(NOISE).NE.A) NOISE = NOISE +1
C
NTOT \(=\) NSIG \({ }^{*}\) IMULMX + NOISE
IF(NTOT.GT.NDIM) STOP 'NTOT IS GREATER THAN NDIM. PROG. ABORTED'

\section*{C \\ C INPUT VARIABLES}

C
1 WRITE \((6,2)\)
2 FORMAT (/,/,/,/,T20,35('*'),/,T20,'*',33X,'*',/,T20,
\& '* DISTRESS SIGNAL DELAY ESTIMATOR *'/,T20, \& '*',33X,'*',/,T20,35('*'),/,/,/,T3,
\& 'C/No of signal \#1 and \#2 (dB-Hz)')
READ (5,*) CNO1,CNO2
C
10 WRITE \((6, *)\) 'Sampling rate multiplication factor ( > 0; < ',IMULMX,')'
READ(5,*) IMUL
IF(IMUL.LE.0) GOTO 10
C
20 WRITE (6,22) DELMX*1.E+3
22 FORMAT (T3,'Delay to be simulated (milliseconds) (must be ',
\& 'less than ',F8.2,' ms )')
READ (5,*) DELAY
DELAY=ABS(DELAY)*1.E-3

\section*{Listing of Program 'Delay estimator.f' (Continued)}
```

            IF(DELAY.GT.DELMX) GOTO 20
    C
c get the file names and open the files to store the result
c summary and the compiled list of errors
C
write (6,*) 'File name where to store summary results'
OPEN (10,FILE=*'File name for summary',STATUS='UNKNOWN')
INQUIRE (10,NAME=FILEN)
WRITE (6,*) FILEN
REWIND (10)
C
write (6,*) 'File name where to store error'
open (20,file=*'File name for error',status='unknown')
rewind (20)
inquire (20,name=filen)
write(6,*) filen
C
c refresh the output window
c
call f_DrawOutpWindow
c
25 write (6,*) 'How many runs do you want?'
read (5,*) NRUNS
if(nruns.le.0) goto 25
c write (6,*) 'Do you want to enter the random number generator seeds (Y or N
read (5,*) ANS
if(ans.eq.'y' .or. ans.eq.'Y') then
WRITE (6,*) 'ENTER SEED1, SEED2'
READ (5,*) SEED1,SEED2
END IF
C
28 WRITE (6,30)
30 FORMAT (/,/,T3,'...computing',/,)
C
C compute the new sampling frequency FS
C Compute the standard deviation of the noise to generate the
C the appropriate C/No's
C Compute the number of samples of delay
C Compute the number of signal samples
C Compute the total number of samples including noise
C Compute the number of samples per bit
C generate the signals. No phase advance for the first signal (ADV=0)
C
FS=FSMPL*FLOAT(IMUL)
CALL SDEV (FSMPL,CNO1,SIGMA1)
CALL SDEV (FSMPL,CNO2,SIGMA2)
C
NDELAY = IFIX((DELAY)*FS+0.1)
NSIGNAL = NSIG*IMUL
A = DELMX*FS
NOISE = IFIX(A)
IF(FLOAT(NOISE).NE.A) NOISE = NOISE + 1
NTOT = NSIGNAL + NOISE

```

\section*{Listing of Program 'Delay estimator.f' (Continued)}

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C

IF(NTOT.GT.NDIM) STOP 'NTOT IS GREATER THAN NDIM. PROG. ABORTED' \(\mathrm{NSAM}=\operatorname{IFLX}(\mathrm{FS} / \mathrm{DRATE} / 2.0+0.1)\)
c
ibegin \(=\) jsecnds \((0)\)
do 1000 ir \(=1\), nruns
\(\mathrm{ADV}=0\).

\section*{CALL ELTLowpass (NSIGNAL,NBIT,NSAM,FS,SIGMA1,SEED1, ADV,FRAME,SIGNAL1)}

\section*{C}

C add noise at the end of the first signal because it is
C assumed that signal\#1 is in advance to signal \#2
C
CALL GAUSSRN (XMEAN,SIGMA1,SEED1,R1,Q1)
IBEEN \(=0\)
DO I \(=\) NSIGNAL +1 ,NTOT
SIGNAL1(I) = R1
IF(IBEEN.EQ.1) THEN
SIGNAL1(I)=Q1
CALL GAUSSRN (XMEAN,SIGMA1,SEED1,R1,Q1) IBEEN \(=0\)
ELSE
IBEEN \(=1\)
END IF
END DO

\section*{C}

C DEFINE SIGNAL NUMBER 2
C Compute required phase advance in terms of fraction of sample
C add noise preceding the signal by the amount of delay
ADV \(=\) DELAY*FS - FLOAT(NDELAY)
\(x \quad\) write \((6, *)\) 'Phase advance in samples = ',ADV
CALL GAUSSRN (XMEAN,SIGMA2,SEED2,R2,Q2)
IBEEN \(=0\)
DO \(I=1\), NDELAY
SIGNAL2(I) = R2
IF (IBEEN.EQ.1) THEN
SIGNAL2(I)=Q2
CALL GAUSSRN (XMEAN,SIGMA2,SEED2,R2,Q2)
IBEEN \(=0\)
ELSE
IBEEN \(=1\)
END IF
ENDDO
C add the signal itself
\[
N=N D E L A Y+1
\]
\&
CALL ELTLowpass (NSIGNAL,NBIT,NSAM,FS,SIGMA2,SEED2,

\section*{Listing of Program 'Delay estimator.f' (Continued)}

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249 x
250 x
251 x
252 x
253
254 C compute cross-correlation
255 C FFT size = NSIZE
256 C
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265 x
266 x
267 x
268 x
269 x
270 x
C
C Fill up rest of vector with noise samples
C
\(\mathbf{N}=\mathbf{N}+\mathbf{N S I G N A L}\)
DO I=N,NTOT
SIGNAL2(I) \(=\) R2

IBEEN \(=0\)
ELSE
\(\operatorname{IBEEN}=1\)
END IF
END DO
C
C signals are now both defined
C
\(x \quad\) READ \((5, *)\) ANS
\(\operatorname{read}(5, *)\) in,is
in=max (in,1)
is \(=\min\) (ntot, is)
is \(=\max (\mathrm{is}, \mathrm{in})\)
write (6,*) FLLEN
REWIND (98)

ENDFILE(98)
CLOSE (98)
END IF
C ENDIF
\&
C
IF (IER.NE.0) THEN

\section*{END IF}

C
read (5,*) i
if(i.eq. 1) then

IF(IBEEN.EQ.1) THEN
SIGNAL2(I)=Q2
CALL GAUSSRN (XMEAN,SIGMA2,SEED2,R2,Q2)
\(x \quad\) write (6,*) 'Save the signal on a file (Y/N)'

IF(ANS.EQ.'Y' .OR. ANS.EQ.'y') THEN
write \((6, *)\) 'from what to what (maximum=',ntot,')'

WRITE \(\left(6,{ }^{*}\right)^{\prime}\) FILE NAME where to save data'
OPEN (98,FILE=*,STATUS='UNKNOWN')
inquire ( 98, name=filen)

WRITE (98,'(I8,A1,F15.7,A1,F15.7)')(I,CHAR(9),SIGNAL1(I)
\& , CHAR(9),SIGNAL2(I),I=in,is)

CALL CORRELATE (SIGNAL1,SIGNAL2,NTOT,LAG,CROSS,NSIZE,

WRITE \((6, *)\) 'ERROR CODE FROM CORRELATE = ',IER
STOP 'PROGRAM ABORTED'
\(x \quad\) write \(\left(6,{ }^{*}\right)\) 'set 1 to store correlation levels'
write ( \(6,{ }^{*}\) ) 'file name where to store correlation levels'
open ( 2, file \(=\) *,status='unknown')
inquire ( 2 ,name=filen)

\section*{Listing of Program 'Delay estimator.f' (Continued)}


\section*{Listing of Program 'Delay estimator.f' (Continued)}

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\section*{368}

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\section*{370}

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200 format(2(2x,f),2x,i,2(2x,f))
itime=jsecnds(ibegin)
if(itime.gt.3600) then
do }\textrm{k}=10,20,1
inquire (unit=k,name=filen)
close (unit=k,status='keep')
open (unit=k,file=filen,status='old',access='append')
end do
ibegin=jsecnds(0)
end if
c
1000 continue
C
endfile(20)
close(20)
C
ENDFILE(10)
CLOSE(10)
STOP 'TOURLOU !'
END
C
C Program to compute the cross-correlation between two signals
C
C Based on :"Programs for Digital Signal Processing",
IEEE Press, 1979, chapter 2.2.
C
C X(),Y() : arrays of N samples to be correlated
CROSS(i): :arrays of LAG values giving the correlation
of }\textrm{X}()\mathrm{ and }\textrm{Y}()\mathrm{ at lag (i-1)
C NSIZE is the FFT size and is related to the required
number of lags using LAG=NSIZE/2+1
C IER is the error code =0 indicates no errors
C
C
=1 means NSIZE is greater than the
maximum size allowed by the program
=2 means the dimension i.e. LAG of CROSS is not
large enough for the requested NSIZE
M. Caron Jan 1990
C
C
C
C
PARAMETER (MXSIZ=4096)
DIMENSION X(N),Y(N),CROSS(LAG)
COMPLEX XX(MXSIZ),XMN,XI,YI,Z(MXSIZ/2+1)
C
C Check for errors
IER =0
IF(NSIZE.GT.MXSIZ) THEN
IER=1
RETURN

```

\title{
Listing of Program 'Delay estimator.f' (Continued)
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\begin{tabular}{l}
C \\
x \\
C \\
\hline
\end{tabular}
C C

ELSE IF (LAG.LT.NSIZE/2+1) THEN
IER=2 RETURN

\section*{END IF}

C
C define variables
C LSHFT = overlap factor per section
C MHLF1 = maximum shift of interest in samples
C NSECT = number of sections of NSIZE samples with overlap of NSIZE/2
C NRD = number of samples to reaf each time
C IW +1 = index of next sample to be read
C NRDY = number of samples to be read on Y()
\(C\) NRDX = number of samples to be read on \(X\) ()
C
```

        LSHFT = NSIZE/2
        MHLF1 = LSHFT+1
        NSECT = FLOAT(N+LSHFT-1)/FLOAT(LSHFT)
        NRD = LSHFT
    ```
        IW \(=0\)
            NRDY = NSIZE
            NRDX \(=\) LSHFT

\section*{C}

C initialize the temporary cross-correlation array
C
DO I=1,MHLF1
\(\mathrm{Z}(\mathrm{I})=\operatorname{CMPLX}(0 ., 0\).
END DO
C
\(x\) WRITE (6,*) 'NSECT = ',NSECT
NSECT1 = NSECT -1
C
C compute the number of sections betwee
C let him know the current status
C
\(\mathrm{x} \quad \mathrm{ND}=(\mathrm{FLOAT}(\mathrm{NSECT}) / 10 .+0.5)\)
C
C
C Compute the FFT of each section of X and Y . Use the odd/even technique
C to compute both over a single FFT for each section
C Accumulate in the frequency domain and then inverse FFT to obtain the
C correlation
C
DO \(190 \mathrm{~K}=1\),NSECT

\section*{C \\ C PRINT K FROM TIME TO TIME TO MONITOR PROGRESS}

C IF THE LAST SECTION, THEN CHECK IF THE NUMBER OF SAMPLES REMAINING C IS EEQUAL TO NSIZE. IF NOT, FILL IN WITH ZEROS.

IF(K.GE.NSECT1) THEN
NRDY \(=\) N \(-(\mathrm{K}-1)^{*}\) LSHFT
IF(K.EQ.NSECT) NRDX \(=\) NRDY

\title{
Listing of Program 'Delay estimator.f' (Continued)
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IF(NRDY.NE.NSIZE) THEN
NRDY1 \(=\) NRDY +1
DO 100 I=NRDY1,NSIZE
\(\mathrm{XX}(\mathrm{I})=\operatorname{CMPLX}(0 ., 0\).
END IF
END IF
C
C read NRDY data starting at sample \#IW+1
C
DO 120 I \(=1\),NRDY
\(\mathrm{J}=(\mathrm{I}+\mathrm{IW})\)
\(\mathrm{XX}(\mathrm{I})=\operatorname{CMPLX}(\mathrm{X}(\mathrm{J}), \mathrm{Y}(\mathrm{J}))\)
C
NRDX1 \(=\) NRDX +1
DO 170 I=NRDX1,NSIZE
\(\mathrm{XX}(\mathrm{I})=\operatorname{CMPLX}(0 ., \operatorname{AIMAG}(\mathrm{XX}(\mathrm{I})))\)
170
C
C correlate X and Y and accumulate \(\operatorname{CONJG}(\mathrm{X}) * \mathrm{Y}\)
C

\section*{CALL FFT (XX,NSIZE,0)}

C
DO 180 I=2,LSHFT
J=NSIZE+2-I
\(\mathrm{XI}=(\mathrm{XX}(\mathrm{I})+\mathrm{CONJG}(\mathrm{XX}(\mathrm{J}))) / 2\).
\(\mathrm{YI}=\mathrm{XX}(\mathrm{J})-\mathrm{CONJG}(\mathrm{XX}(\mathrm{I})) / 2\).
\(\mathrm{YI}=\operatorname{CMPLX}(\operatorname{AIMAG}(\mathrm{YI}), \mathrm{REAL}(\mathrm{YI}))\)
\(\mathrm{Z}(\mathrm{I})=\mathrm{Z}(\mathrm{I})+\mathrm{CONJG}(\mathrm{XI})^{*} \mathrm{YI}\)
180
C
\[
\mathrm{XI}=\mathrm{XX}(1)
\]
\[
\mathrm{Z}(1)=\mathrm{Z}(1)+\mathrm{CMPLX}(\operatorname{REAL}(\mathrm{XI}) * \operatorname{AIMAG}(\mathrm{XI}), 0 .)
\]
\(\mathrm{XI}=\mathrm{XX}(\mathrm{MHLF} 1)\)
\(\mathrm{Z}(\mathrm{MHLF} 1)=\mathrm{Z}(\mathrm{MHLF} 1)+\mathrm{CMPLX}(\mathrm{REAL}(\mathrm{XI}) * \operatorname{AIMAG}(\mathrm{XI}), 0\).
IW \(=\mathrm{IW}+\mathrm{LSHFT}\)
190 CONTINUE
C COMPUTE INVERSE DFT FOR CORRELATION
C
DO 200 I \(=2\), LSHFT
J=NSIZE+2-I
\(\mathrm{XX}(\mathrm{I})=\mathrm{Z}(\mathrm{I})\)
\(\mathrm{XX}(\mathrm{J})=\operatorname{CONJG}(Z(\mathrm{I}))\)
200 CONTINUE
C
\(\mathrm{XX}(1)=\mathrm{Z}(1)\)
\(\mathrm{XX}(\mathrm{MHLF} 1)=\mathrm{Z}(\) MHLF1 \()\)
C

> CALL FFT(XX,NSIZE,1)

\section*{C}

C normalize the results

\section*{Listing of Program 'Delay estimator.f' (Continued)}

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C
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        FN = FLOAT(N)
        DO I=1,MHLF1
                CROSS(I) = REAL(XX(I))/FN
        END DO
        RETURN
        END
    C
C Program to compute the DFT AND IDFT
C Based on :"Programs for Digital Signal Processing",
IEEE Press, 1979, chapter 2.2.
C INV = 0 ==> DFT, INV =1 ->> IDFT
C C
C X(): COMPLEX ARRAY OF N NUMBERS
C
C-------------------------------------------------------
SUBROUTINE FFT (X,N,INV)
COMPLEX X(N),U,W,T
DATA PI/3.14159265358979324/
C
M=[FIX(ALOG(FLOAT(N))/ALOG(2.)+0.1)
NV2 = N/2
NM1 = N-1
J=1
DO }40\mathrm{ I= 1,NM1
IF(I.LT.J) THEN
T=X(J)
X(J)=X(I)
X(I)=T
END IF
K=NV2
IF(K.LT.J) THEN
J=J-K
K=K/2
GOTO 20
ENDD IF
J=J+K
30
40 CONTINUE
DO 70 L=1,M
LE=2**L
LE1 = LE/2
U = CMPLX (1.0.)
W = CMPLX(COS(PI/FLOAT(LE1)),-SIN(PI/FLOAT(LE1)))
IF(INV.NE.0) W=CONJG(W)
DO }60\textrm{J}=1,\textrm{LE
DO }50\mathrm{ I=J,N,LE
IP=I+LE1
T=X(IP)*U
X(IP)=X(I)-T

```

\title{
Listing of Program 'Delay estimator.f' (Continued)
}
\begin{tabular}{lll}
541 & & \multicolumn{1}{c}{\(\mathrm{X}(\mathrm{I})=\mathrm{X}(\mathrm{I})+\mathrm{T}\)} \\
542 & 50 & \multicolumn{1}{c}{ CONTINUE } \\
543 & & \multicolumn{1}{c}{ U=U*W } \\
544 & 60 & CONTINUE \\
545 & 70 & CONTINUE \\
546 & & IF(INV.EQ.0) RETURN \\
547 & & DO 80 I=1,N \\
548 & 80 & X(I) \(=\mathrm{X}(\mathrm{I}) /\) CMPLX(FLOAT(N),0.) \\
549 & & RETURN \\
550 & & END
\end{tabular}

\section*{Listing of Subroutine 'EltLowpass.f'}
```

1
SUBROUTINE EITLOwpass (NSIG,NBIT,NSMPI,FSMPI,SIGMAI,SEED, $\&$
ADV, FRAME, SIGNAL)
3
4 C This program generates the samples of a simulated ELT/EPIRB
5 beacon.
6 C A sampling rate of FSMPL Hz is used.
8C
fc = nominal carrier frequency
INTEGER FRAME (NBIT) , MESSAGE (112)
REAL SIGNAL(1), INC
REAL*4 R, Q, SIGMA, XMEAN
INTEGER*4 SEED
DATA XMEAN /0.0/
DATA PI/3.14159265358979324/
DATA MESSAGE $/ 1,1,1,1,1,1,1,1,1,1,1$
$\&_{\&} \quad, 1,1,1,1,0,0,0,1,0,1,1$
\& $\quad 1,1,0,1,1,0,1,0,1,1,1$
\& $, 1,0,0,1,1,1,0,0,0,0,1$
\& $\quad, 1,1,0,0,0,0,0,1,0,0,1$

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\section*{Listing of Subroutine 'EltLowpass.f' (Continued)}

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\& \(\quad, 1,1,0,0,1,0,1,0,0,0,0\)
\& \(\quad 0,1,0,0,0,1,0,0,0,1,0\)
\& \(\quad, 1,0,0,0,0,0,1,0,0,0,1\)
\& \(, 1,1,0,0,1,0,1,1,0,0,0\)
\& \(\quad 0,1,1,1,0,0,0,0,0,1,0\)
\& , 1,0 /

C
```

        PIBY4=PI/4.
    ```
        TWOP \(I=2.0 * P I\)
C
C A frame of ELT signal will be simulated. First, read from
file
\(C\) FRAME.DAT the precomputed data bits and store them in array
FRAME.
C
c OPEN(UNIT=99,FILE='FRAME.DAT',STATUS='OLD')
c \(\operatorname{READ}(99, *)\) ( \(\operatorname{FRAME}(I), I=1, \operatorname{NBIT})\)
c CLOSE (UNIT=99,STATUS='KEEP')
do \(i=1, n b i t\)
            frame (i)=message (I)
end do
C
\(C\) SUM and No are used to compute the generated standard deviation
\(C\) and to compare it with the requested one at the end
SIGMA \(=\) SIGMAI
\(\mathrm{x} \quad \operatorname{SUM}=0\).
\(x \quad\) NO \(=0\)
\(\mathrm{x} \quad\) SUM2 \(=0\).
ISMPL+1 : next sample index for array SIGNAL
C OMEGA \(=2 \star\) pi*frequency
C compute the phase increment per sample
\(C\) compute the number of samples for the pure carrier
C preceding the message
C compute random phase shift between 0 and 2 pi (constant for
burst)
\(C\) set first phase value (PARAM) such that sample starts
\(C\) with ADV*INC plus random phase shift
C
ISMP L=0
NCARR \(=\operatorname{IFIX}(160 . E-3 * F S M P L+0.1)\)
C \(\quad \mathrm{R}=0\).
c \(\quad Q=T W O P I\)
c CALL UNIFORM ( \(R, Q, \operatorname{SEED}, Q\) )
c \(\quad\) PARAM \(=Q\)
param \(=0\).
C
Now, generate \(160 \mathrm{~ms}\left(+\right.\) or \(-1 \%\) ) (or \(160 \cdot \mathrm{E}-3 * \mathrm{ESMPL}=\mathrm{N}^{\left(\mathrm{A}^{\mathrm{R}}\right)}\)
C sample intervals.
PARAM \(=\) AMOD (PARAM, TWOPI)
\(S S=\sin (\) param + piby 4\()\)
DO ICARR=1, NCARR, 2

\section*{Listing of Subroutine 'EltLowpass.f' (Continued)}

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C
C
C
                    ISMP L=ISMPL +1
                CALL GAUSSRN (XMEAN, SIGMA, SEED, R, Q)
                SIGNAL \((I S M P L)=S S+R\)
                    ISMPL=ISMPL+1
                SIGNAL \((I S M P L)=S S+Q\)
                    \(\mathrm{x} \quad\) SUM2 \(=\operatorname{SUM} 2+\mathrm{R} * * 2+\mathrm{Q} * * 2\)
                    \(x \quad S U M=S U M+R+Q\)
                \(\mathrm{NO}=\mathrm{NO}+2\)
            END DO
            ISMPL \(=\) NCARR
            C
            c
c
c generator produces two random numbers per call
C IC=0 if NSMPI is even IC=1 otherwise
    \(I C=1\)
    \(I F((N S M P L / 2) * 2 \cdot E Q \cdot N S M P L) \quad I C=0\)
    There will be a nominal (FSMPL/800) = NSMPL samples per phase
        of the Manchester symbol where the data rate is 400 bps
    Now modulate the carrier with the frame data bits
            \(S S=P A R A M+P I B Y 4\)
            DO IBIT=1, NBIT
            IF (FRAME (IBIT).EQ.1) THEN
\(C\) when bit is 1 , modulation is \(+1.1,-1.1\) radians
                    DO JSMPL=1, NSMPL, 2
                                    ISMPI \(=1\) SMP +1
                                    CALL GAUSSRN (XMEAN,SIGMA, SEED, R, Q)
                                    SIGNAL (ISMPL) \(=\) SIN (1.1+SS) + R
                    ISMPL=ISMPL+1
                                SIGNAL \((\) ISMPL \()=S I N(S S+1.1)+Q\)
                                SUM2 \(=\) SUM2 \(+\mathrm{R}^{* * 2+Q * * 2 ~}\)
                                \(S U M=S U M+R+Q\)
                                \(\mathrm{NO}=\mathrm{NO}+2\)
                                END DO
                    \(I S M P I=I S M P L-I C\)
                    DO JSMPL=1, NSMPL, 2
                                    ISMP L=ISMP +1
                                    CALL GAUSSRN (XMEAN, SIGMA, SEED, R, Q)
                                    SIGNAL (ISMPL) \(=\) SIN \((S S-1.1)+\) R
                                    \(\operatorname{ISMPL}=I S M P L+1\)
                                    SIGNAL \((I S M P I)=S I N(S S-1.1)+Q\)

                                    SUM \(=\) SUM \(+R+Q\)
                                    \(\mathrm{NO}=\mathrm{NO}+2\)
        END DO
                \(I S M P L=I S M P L-I C\)
            ELSE

\section*{Listing of Subroutine 'EltLowpass.f' (Continued)}

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172 x
173 x
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183 x
184 x
185 x
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195 x
SMEAN \(=\) SUM/FLOAT (NO)
196 x STDEV=SQRT (SUM2/FLOAT (NO) - SMEAN**2)
197 C
198
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200
\(201 \times \quad \& \quad\) F15.7,/,T5,'Standard Deviation Obtáned \(=1, F 15.7,1, T^{51}\) \(202 \mathrm{x} \& \quad\) 'Absolute error \((\%)=1, \mathrm{~F} 7.2, /, /\) )
203 c WRITE \((10,100)\) SMEAN,SIGMA,'STDEV', ABS (SIGMA-STDEV)/SIGMA
204 C
204
205 206
\(C\) when the bit is a zero, the modulation is \(-1.1,+1.1\) radians C

DO JSMPL=1,NSMPL,2
ISMPL=ISMPL+1
CALL GAUSSRN (XMEAN, SIGMA, SEED, R, \(Q\) )
SIGNAL (ISMPL) \(=\) SIN (SS-1.1) + R
ISMPL=ISMPL+1
SIGNAL \((\) ISMPL \()=\) SIN \((S S-1.1)+Q\)
SUM2 \(=\) SUM2 \(+\mathrm{R}^{\star * 2}+\mathrm{Q}^{\star *} 2\)
SUM \(=\operatorname{SUM}+\mathrm{R}+\mathrm{Q}\)
\(\mathrm{NO}=\mathrm{NO}+2\)
END DO
ISMPL \(=\) ISMPL - IC
DO JSMPL=1,NSMPL,2
ISMP \(L=I S M P L+1\)
CALL GAUSSRN (XMEAN, SIGMA, SEED, R, Q)
SIGNAL (ISMPL) =SIN (SS+1.1) +R
ISMP \(L=I S M P L+1\)
SIGNAL (ISMPL) \(=\) SIN \((S S+1.1)+Q\)
SUM2 \(=\) SUM2 \(+R^{\star *} 2+Q^{* * 2}\)
SUM \(=\) SUM \(+R+Q\)
\(\mathrm{NO}=\mathrm{NO}+2\)
END DO
ISMPL = ISMPL - IC
ENDIF
END DO
C
C compute the standard deviation produced by the random number
C generator and write the absolute error relative to the
requested one
C
x WRITE \((6,100)\) SMEAN,SIGMA,STDEV,ABS (SIGMA-STDEV)/SIGMA*10 100 FORMAT (/,T3,'FROM ELTSIG',/,T5,' Mean = ', F15.7, 1 , \& 'Required standard Deviation \(=\) ',
\(\operatorname{WRITE}(10,100)\) SMEAN,SIGMA,STDEV,ABS (SIGMA-STDEV)/SIGMA*100
RETURN
END

\section*{Listing of File 'Random Number.f'}
\begin{tabular}{|c|c|c|}
\hline 1 & & \\
\hline 2 & & C \\
\hline 3 & & \(C\) Subroutine to generate a random number between \\
\hline 4 & & \(C A\) and \(B\) with uniform distribution. \\
\hline 5 & & C \\
\hline 6 & C & \(C\) ISEED is the seed to be used at input and is \\
\hline 7 & & \(C\) the next seed for the next call at output \\
\hline 8 & & \(C\) RAND is the random number \\
\hline 9 & C & C \\
\hline 10 & C & C It is suggested to use an initial seed of 4 for \\
\hline 11 & C & the first call. \\
\hline 12 & C & \\
\hline 13 & C & Ref. "Mathematical Methods for Digital Computers Vol II', \\
\hline 14 & C & John Wiley and Sons, Inc., 1967, Chap. 12 \\
\hline 15 & C & John Wiley and Sons, Inc., 1967, Chap.12 \\
\hline 16 & C & M. Caron \\
\hline 17 & c & Jan 1990 \\
\hline 18 & C & \\
\hline 19 & C & \\
\hline 20 & C & \\
\hline 21 & & SUBROUTINE UNIFORM ( \(\mathrm{A}, \mathrm{B}\), ISEED, RAND) \\
\hline 22 & & REAL* 4 A, B, RAND \\
\hline 23 & & INTEGER*4 ISEED \\
\hline 24 & C & \\
\hline 25 & C & ISEED \(=\) ISEED*65539 \\
\hline 26 & C & RAND \(=\) FLOAT (IABS (ISEED) \() /(2.0 * * 31)\) \\
\hline 27 & c & \\
\hline 28 & C & use fortran routine to generate the number \\
\hline 28
30 & C & 俍 \\
\hline 30 & & RAND \(=\) RAN (ISEED) \\
\hline 31 & & RAND \(=\) RAND* \((\mathrm{B}-\mathrm{A})+\mathrm{A}\) \\
\hline 32 & & RETURN \\
\hline 33 & & END \\
\hline 34 & &  \\
\hline 35 & C & \\
\hline 36 & C & Subroutine to generate a pair of random numbers with \\
\hline 37 & C & Gaussian distribution with mean XMEAN and standard \\
\hline 38 & & deviation DEV. \\
\hline 39 & C & \\
\hline 40 & C & ISEED : is the seed to be used and the new one is \\
\hline 41 & C & returned \\
\hline 42 & c & DEV : is the standard deviation \\
\hline 43 & C & XMEAN : mean \\
\hline 44 & c & R : is the first random number \\
\hline 45 & C & Q : is the second random number \\
\hline 46 & C & \\
\hline 47 & & It is suggested to used ISEED=4 for the first call \\
\hline 48 & C & \\
\hline 49 & C & Ref. "Mathematical Methods for Digital Computers Vol II', \\
\hline 50 & C & Re. John Wiley and Sons, Inc., 1967, Chap. 12 \\
\hline 51 c & C & \\
\hline 52 c & c & \\
\hline 53 & C & \\
\hline 54 c & & M. Caron \\
\hline
\end{tabular}

\section*{Listing of File 'Random Number.f' (Continued)}
\begin{tabular}{|c|c|c|}
\hline 55 & C & Jan 1990 \\
\hline 56 & C & \\
\hline 57 & c & \\
\hline 58 & c & \\
\hline 59 & & SUBROUTINE GAUSSRN (XMEAN, DEV, ISEED, R, Q) \\
\hline 60 & C & \\
\hline 61 & & INTEGER*4 ISEED \\
\hline 62 & & REAL*4 DEV, XMEAN,R,Q \\
\hline 63 & & DATA TWOPI/6.28318530717959/ \\
\hline 64 & c & \\
\hline 65 & C & \(P I=3.14159265358979324\) TWOPI \(=2.0 * P I\) \\
\hline 66 & C & \\
\hline 67 & c & CALL UNIFORM (0.,1., ISEED,R) \\
\hline 68 & c & CALL UNIFORM (0.,1.,ISEED,Q) \\
\hline 69 & c & \\
\hline 70 & c & use Fortran routine to generate the uniformly \\
\hline 71 & C & distributed numbers \\
\hline 72 & c & \\
\hline 73 & & R=ran (iseed) \\
\hline 74 & & Q \(=\) ran (iseed) \\
\hline 75 & & \(\mathrm{S}=\mathrm{SQRT}(-2.0 *\) ALOG (R) ) \\
\hline 76 & & TWORIQ \(=\) TWOPI*Q \\
\hline 77 & & \(\mathrm{R}=\mathrm{S}\) COS (TWOPIQ) * DEV + XMEAN \\
\hline 78 & & \(Q=S * S I N(T W O P I Q) * D E V ~+~ X M E A N ~\) \\
\hline 79 & & RETURN \\
\hline 80 & & END \\
\hline
\end{tabular}

\section*{Listing of File 'SDEV.f'}


\section*{Listing of Script File to Compile and Link}
```


# File: 'Delay Estimator.make'

    Target: 'Delay Estimator'
    Sources: 'Delay Estimator.f' ELTLowpass.F 'random number.'
    SDEV.F

# Created: Friday, September 14, 1990 8:24:27 AM

```
```

'Delay Estimator' ff 'Delay Estimator.make' {OBJECTS}

```
'Delay Estimator' ff 'Delay Estimator.make' {OBJECTS}
    Link -f -srt -ad 4 -ss 1000000 -w -t APPL -c '????' \partial
    Link -f -srt -ad 4 -ss 1000000 -w -t APPL -c '????' \partial
        {OBJECTS} \partial
        {OBJECTS} \partial
        "{Libraries}"Runtime.o \partial
        "{Libraries}"Runtime.o \partial
        "{Libraries}"Interface.o \partial
        "{Libraries}"Interface.o \partial
        "{FLibraries)"FORTRANlib.O \partial
        "{FLibraries)"FORTRANlib.O \partial
        "{FLibraries}"IntrinsicLibFPU.o \partial
        "{FLibraries}"IntrinsicLibFPU.o \partial
        "{FLibraries}"FSANELibFPU.O \partial
        "{FLibraries}"FSANELibFPU.O \partial
        -O 'Delay Estimator'
        -O 'Delay Estimator'
    Delay Estimator.f.O' f 'Delay Estimator.make' 'Delay Estim'
    Delay Estimator.f.O' f 'Delay Estimator.make' 'Delay Estim'
    FORTRAN -mc68020 -mc68881 -opt=3 'Delay Estimator.f'
    FORTRAN -mc68020 -mc68881 -opt=3 'Delay Estimator.f'
ELTLowpass.F.O f 'Delay Estimator.make' ELTLowpass.F
ELTLowpass.F.O f 'Delay Estimator.make' ELTLowpass.F
    FORTRAN -mc68020 -mc68881 -opt=3 ELTLowpass.F
    FORTRAN -mc68020 -mc68881 -opt=3 ELTLowpass.F
'random number.f.o' f 'Delay Estimator.make' 'random number.f
'random number.f.o' f 'Delay Estimator.make' 'random number.f
    FORTRAN -mc68020 -mc68881 -opt=3 'random number.f'
    FORTRAN -mc68020 -mc68881 -opt=3 'random number.f'
SDEV.F.o f 'Delay Estimator.make' SDEV.F
SDEV.F.o f 'Delay Estimator.make' SDEV.F
    FORTRAN -mc68020 -mc68881 -opt=3 SDEV.F
```

    FORTRAN -mc68020 -mc68881 -opt=3 SDEV.F
    ```

\section*{Run Example}
```

*********************************************

```

```

* DISTRESS SIGNAL DELAY ESTIMATOR *
*     * 

****************************************
C/No of signal \#1 and \#2 (dB-Hz)
30,30
Sampling rate multiplication factor (>0; < 52)
26
Delay to be simulated (milliseconds) (must be less than 18.00 ms )
5
File name to store summary results
example.dat
File name where to store error
error.dat
How many runs do you want?
1
Do you want to enter the random number generator seeds (Y or N)
N
...computing
C/No of signal \#1,\#2 = 30.00 30.00(dB-Hz)
Simulated Delay = 5.00(ms)= 208 (samples)
Maximum Cross-Correlation = 0.4816 and
occured at samples \# 208 corresponding to a delay of }\quad5.0000000\textrm{ms
Delay error = 0.000\mus Next seeds (1 and 2 ) = 1548668310 1015696458
STOP TOURLOU!

```

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This report looks at techniques to perform the localization of SARSAT 406 MHz distress beacons using geostationary satellites. After reviewing the characteristics of the distress beacon transmitters, the main specifications of the GOES satellites are described as they can be assumed as typical geostationary satellites for this application. A brief review of the geolocation techniques is presented and an in-depth analysis is conducted on the time difference of arrival technique. The analysis covers the impact of the Earth's flatteness, channel impairments and satellite geometry. The delay estimator performance is expressed as a function of signal-tonoise ratio and circular error probability. The overall system performance is directly related to the performance of the delay estimator.

Results of both theoretical analysis and computer simulations show that the accuracy of the position derived from the time difference of arrival technique is limited by (1) the low SNR of the distress beacons when relayed via geostationary satellites and, (2) by the low transmission rate (and thus bandwidth) of the beacon signal creating "broad" autocorrelation peaks and making the delay estimation process difficult. In the worst case \(\mathrm{C} / \mathrm{No}\) of \(30 \mathrm{~dB}-\mathrm{Hz}\) and with three satellites spaced at \(30^{\circ}\), we can say with a \(90 \%\) confidence level that the circular error probability is between 3.3 and 4.9 km for a beacon in Canada. This is equivalent to an error not exceeding between 8 and 11.8 km for \(95 \%\) of the time which is comparable to the current system based on a maximum error of 5 km for \(90 \%\) of the time.
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Geolocation
Position location
Geostationary satellite
SARSAT/COSPAS
Satellite Search and Rescue
Time difference of arrival
GOES satellite
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[^0]:    An ellipsoid is an exact Earth cantly the geometry and introduces insignificant errors.

[^1]:    Table 6.2.1-1 Distance Between Position Fixes Generated Using Satellites at $135^{\circ} \mathrm{W}$ and $75^{\circ} \mathrm{W}$ with Range Errors of 0 km and 15 km over a Nominal Range Difference of 200 km .

